# First look to the CONUS+ data

### On behalf of the CONUS Collaboration

MAX-PLANCK-INSTITUT FÜR KERNPHYSIK HEIDELBERG (MPIK)



Applied Antineutrino Physics (Aachen), October 2024

### KKL nuclear power plant

- CONUS+ experiment is operating at the KKL power plant (Leibstadt, Switzerland) since November 2023.
- BWR with high duty-cycle: 1 month/year of reactor-off.
- CONUS+ is placed inside the reactor building in the ZA28R027 room.
- 20.7 m from 3.6 GWt reactor core  $\rightarrow$  high antineutrino flux expected 1.45 x 10<sup>13</sup> V<sub>s</sub> s<sup>-1</sup> cm<sup>-2</sup>
- Reactor dome 1.2 m steel reinforced concrete + 3.8 cm steel containment structure.
- Concrete ZA28R027 room 0.35 m. Average overburden between 7-8 m w.e.



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### CONUS+ location: ZA28R027 room

- Isolated area for CONUS+ with metallic wall. Temperature in room kept stable with AC system.
- Direct network connection to MPIK. <u>Monitoring in</u> real time possible!
- Reactor drywell head over room during reactor off. Thickness 3.8 cm steel → overburden variation 0.25 m w.e.







### **Background characterization campaign**

Extensive background characterization campaign between 2022-2023 at different positions:

- Y measurements with HPGe detector.
- Neutron measurements with Bonner Sphere array system.
- Environmental parameters (radon, temperature ...).
- Cosmic muons with liquid scintillator.
- Vibrations with piezoelectric sensors.
- Surface contamination with wipe tests + low background Ge spectrometers.

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#### **Publication in preparation!!**

## CONUS+ background: **Y**'s

- Ultra-low background p-type coaxial HPGe detector CONRAD (m =2.2 kg). Electrical cryocooling system.
- Scan over different positions with measurement from few hours to one day.
- High energy gamma contribution (>2.7 MeV) factor 25 smaller than at Brokdorf power plant. Stronger contribution of <sup>60</sup>Co lines.





### **CONUS+ background: Cosmic muons**

- Liquid scintillator cell filled with 120 ml of "Ultima Gold". PMT for light detection.
- Measurements at MPIK and KKL during off time for comparison.
- Quality cuts applied: saturation, pile-up.
- Pulse shape discrimination cut to remove neutrons.





- Energy cut at 3 MeV to avoid environmental radioactivity. Reactor OFF to avoid high energy γ contribution.
- <u>Muon rate surface: 200±5 counts/s/m<sup>2</sup>.</u>
- <u>Muon rate ZA28R027: 107±3 counts/s/m<sup>2</sup>.</u>
- Reduction factor of 1.9 in KKL compared to surface  $\rightarrow$  overburden 7.4 m w.e.
- Muon rate factor 2.4 larger than at KBR. Overburden at KBR 24 m w.e.
- Impact reactor drywell head 0.25 m w.e.

### **CONUS+** background: Radon

- Radon can diffuse into the detector chamber and produce some background.
- Monitoring of the radon level in the room during one year.





- Radon concentration average value of 110 Bq/m<sup>3</sup>. Significant fluctuations.
- Flushing with bottles filled with air and stored for periods over 3 weeks.
- Background reduction in [100-400] keV range by factor 5. Radon lines (242, 295 and 352 keV) strongly suppressed.

### **CONUS+** background: Reactor neutrons

- Neutron spectrometry with Bonner Sphere detectors in scientific cooperation with PSI.
- Monitoring neutron rate with 5" PE sphere. Correlation with thermal power. Most neutrons in the room are produced by the reactor.
- Same configuration of spheres as in KBR for direct comparison giving a sensitivity from 10<sup>-9</sup> to 10<sup>3</sup> MeV



### **CONUS+** background: Reactor neutrons

- Measurement in same position with 1 sphere at the time. Neutron flux stable within 3%.
- Unfolded spectra in lethargy representation during reactor on. 80% of the neutrons have energies below 0.4 eV.
- Total neutron flux 284 n/GW/cm<sup>2</sup>/h  $\rightarrow$  37 times larger than in KBR.
- However, simulations show a negligible impact!!



Preliminary Energy region	φ [cm <sup>-2</sup> (GWh) <sup>-1</sup> ]	
Thermal (<0.4 eV)	206	
Intermediate (>0.4 eV && < 0.1 MeV)	74	
Fast (>0.1 MeV && < 19.4 MeV)	4	
Total	284	



### **CONUS+** background: Cosmic neutrons

- Measurement with BSS not conclusive. Count rates 6 times larger than in KBR.
- Alternative approach based on simulations.
- Initial neutron spectra from [1], neutron flux value considered 0.013 n/s/cm<sup>2</sup>.
- Neutrons propagated over reactor building. Neutrons suppressed almost two orders of magnitude, but still large impact over the CONUS+ background.
- Rate variation for cascade neutrons with the reactor drywell head ~16%.
- Muon-induced neutrons in concrete yet not included in this simulation.



[1] P. Goldhagen, J. M. Clem, J. W. Wilson, *Radiation Protection Dosimetry*, Volume 110, Issue 1-4, 1 August 2004, Pages 387–392, https://doi.org/10.1093/rpd/nch216

### **CONUS+** modifications

- 4 refurbished p-type point contact HPGe with total crystal/active mass: 4 kg /3.74kg as target. Better trigger efficiency and threshold.
- Active + passive shielding: low <sup>210</sup>Pb lead, borated and pure PE and 2 active µ-vetos (plastic scintillator).
- Less  $\gamma$ 's and more muons  $\rightarrow$  Third lead layer replaced by additional second muon veto.
- Apply muon veto offline. New DAQ for veto system. Energy deposited stored for each PMT.
- New stainless steel frame to meet KKL earthquake safety requirements.



CONUS Collaboration, arXiv:2407.11912

## Stability during first year @KKL

0.03

0.02

0.01

-0.02

-0.03

×10

1.16

1.14

1.12

1.09

#### Trigger efficiency parameters

Energy calibration



Seasonal variation veto rate. Anticorrelation with temperature (-0.13% /C) and pressure (-0.1% /mbar).

Variation during reactor off. Two effects:

- Reactor correlated high energy γ's from neutron capture.
- Reactor drywell head over room.

## Stability during first year @KKL





Preliminary background rate in [0.4-1.0] keV region 35-45 cts/day (detector dependent).

Reactor neutrons negligible. <sup>210</sup>Pb strongly suppressed.

Dominant component from cosmic rays (muon-induced + cosmogenic neutrons). Still missing background at low energy.

### **Threshold determination**

Energy threshold defined independently for each detector. Two conditions:

- Trigger efficiency over 20% (down to 70 eV<sub>ee</sub>).
- Noise peak contribution below 10% expected CEvNS signal.

Non-linearity at low energies due to DAQ energy reconstruction limitations and lost of trigger efficiency.

Estimated with pulser scan from 2 keV down to few eV. Maximum deviation from linearity 15 eV.

Energy threshold for after non-linearity correction 150-160 eV



# **CEvNS** signal

- Evolution most relevant actinides during typical reactor cycle at KKL.
- Average trigger efficiency during run-1 considered for each detector.
- Lindhard quenching factor with k=0.162.
- With 160 eV energy threshold, 500 events expected in 1 year kg.





# Summary

- Large differences in the background conditions compare to KBR: smaller overburden, less high energy y-rays and more neutrons. Modifications to CONUS+ shield.
- Background dominated by muon-induced and cosmic neutrons.
   Background rate in [0.4-1] keV region ~2 times larger than in CONUS.
- First year of operation with the CONUS+ detector. Stable environmental and detector conditions.
- Energy threshold at KKL 150-160 eV<sub>ee</sub>. Expected CEvNS signal with 1 year kg of data  $\rightarrow$  > 500 events.

### Stay tuned for new results!!



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### Thank you for your attention



### **Surface contamination**

- Larger surface contamination respect to KBR. e.g. <sup>60</sup>Co 300 vs 1200 cts/day/kg.
- Wipe test from "hot spot" will increase CONUS+ background by 5 times.
- Strict cleaning protocols mandatory during installation.



### CONUS+ background: **Y**'s

- Lines produced by <sup>16</sup>N strongly reduced compared to KBR. <u>Larger distance to</u> <u>reactor cooling system</u>.
- Lines from <sup>28</sup>Si and <sup>40</sup>Ca visible at KKL.
   Portland cement with high content of these isotopes.
- Larger contribution from <sup>56</sup>Fe at KKL.
   Larger neutron fluence and more material.
- Larger contribution of <sup>63</sup>Cu from CONRAD cryostat. <u>Larger neutron</u> <u>fluence at KKL.</u>

PC / Energy [keV] /BR[%]	KKL: Ex-HPU-B	KKL: ZA28R027	KBR: ZA408
${ m ^{53}Fe}({\rm n},\gamma){ m ^{54}Fe}$	(reactor structure)		
8787 SEP	$573 \pm 45$	$18.5 \pm 1.5$	$9.6 \pm 0.6$
9298 (100%)	$707 \pm 56$	$19.7 \pm 1.6$	$11.3 \pm 0.5$
${\rm ^{56}Fe}({\rm n},\gamma){\rm ^{57}Fe}$	(reactor structure)		
4217 (23.0%)	$1896{\pm}149$	$78.5{\pm}6.2$	not visible
5920 (33.8%)	$2504 \pm 190$	$95.1 \pm 7.5$	not visible
6018 (34.8%)	$2787 \pm 220$	$98.0 \pm 7.4$	not visible
7120 SEP	6974±596*	$285 \pm 23^*$	not visible
7135 SEP	double peak	double peak	double peak
7278 (20.7%)	$1544 \pm 122$	$80.5 \pm 6.4$	$11.5 \pm 1.1$
7631 (100%)	8717±735*	$363 \pm 29^{*}$	$137 \pm 4^{*}$
7646 (86.2%)	double peak	double peak	double peak
${}^{63}\mathrm{Cu}(\mathrm{n},\gamma){}^{64}\mathrm{Cu}$	(HPGe cryostat)		
7406 SEP	$1992 \pm 157$	$228 \pm 18$	$27.2 \pm 1.3$
7638 (48.9%)	$995 \pm 79^{*}$	$120 \pm 9^*$	$15.7 \pm 1.6^*$
7916 (100%)	$2034 \pm 161$	$245 \pm 19$	$29.7 \pm 1.0$
$^{28}$ Si(n, $\gamma$ ) $^{29}$ Si	(concrete CONUS+ room)		
3539 (100%)	not visible	$276 \pm 22$	not visible
4934 (93.3%)	not visible	$213 \pm 17$	not visible
6379 (16.0%)	not visible	$19.9 \pm 1.6$	not visible
7199 (10.0%)	not visible	$6.8 \pm 0.5$	not visible
${\rm ^{40}Ca}(n,\gamma){\rm ^{41}Ca}$	(concrete CONUS+ room)		
4418 (17.1%)	not visible	$105 \pm 8$	not visible
6419 (43.5%)	not visible	$181 \pm 14$	not visible
${}^{16}O(n,p\gamma){}^{16}N$	(reactor cooling system)		
5617 SEP	$7143 \pm 564$	not visible	$49972 \pm 575$
6128 (67%)	$12652 \pm 998$	not visible	$85086 \pm 979$
7115 (4.9%)	$2526 \pm 199$	not visible	$10097 \pm 116$

### Stability during first year @KKL

Resolution 10.4 keV line



### **Trigger efficiency @KKL**



$$\varepsilon_{trig} = 0.5 \cdot \left( 1 + \operatorname{erf}\left(\frac{E_{ee} - t_1}{t_2}\right) \right)$$

## Energy spectra @KKL

- Veto window 450 µs. Conservative approach as in CONUS. Dead-time ~12%
- Background rejection with new veto system 99% in the [0.4-1] keV region.
- After veto, X-ray lines induced by cosmogenic events in Ge visible. Energy calibration with K and L shells and propagated to low energies. <sup>252</sup>Cf irradiation to improve energy calibration uncertainty <5 eV.</li>

