



# Applied Antineutrino Physics Workshop 2024

28th - 30th October 2024 - Aachen, Germany

## The JUNO and TAO detectors and their physics potential

Davide Chiesa on behalf of the JUNO collaboration

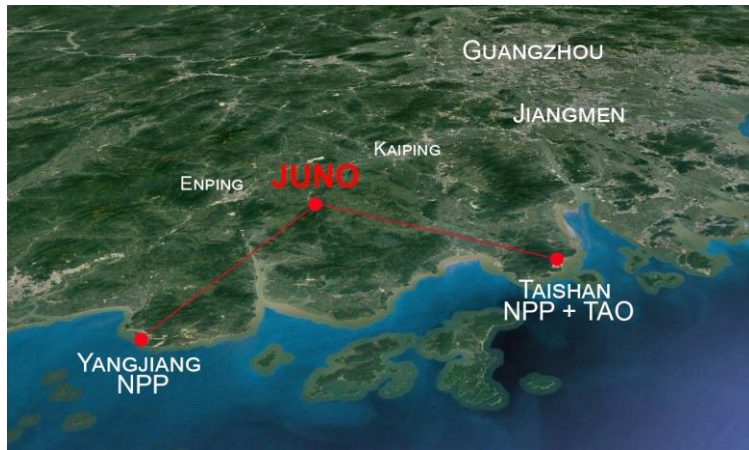
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# The JUNO experiment

- ▶ JUNO (Jiangmen Underground Neutrino Observatory) is a 20 kton liquid scintillator detector located  $\sim 650$  m underground at  $\sim 52.5$  km from two Nuclear Power Plants in China.



- ▶ Construction of the JUNO detector is nearly complete and we expect to start data taking next year.



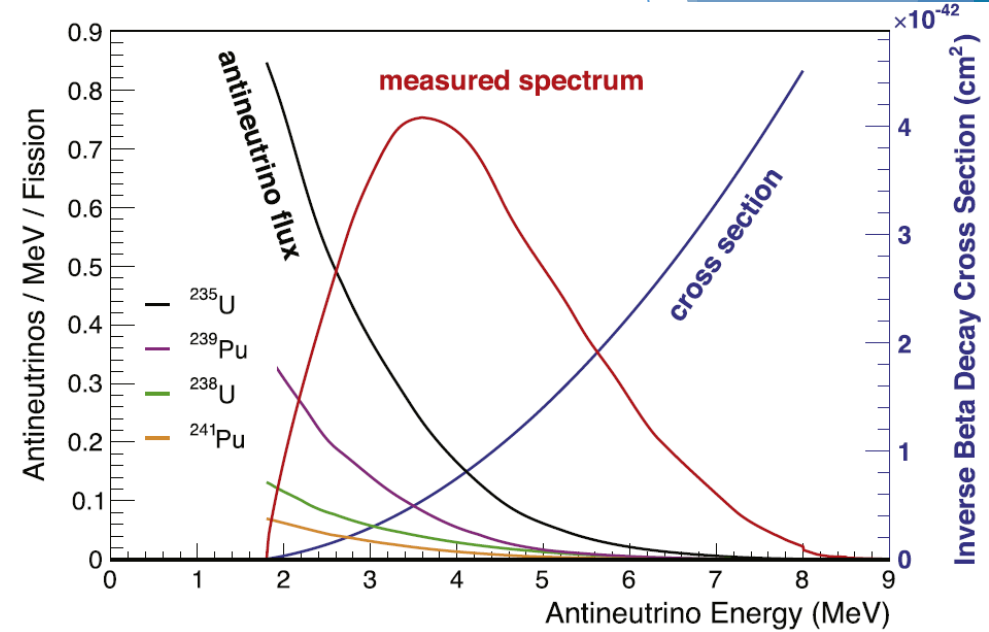
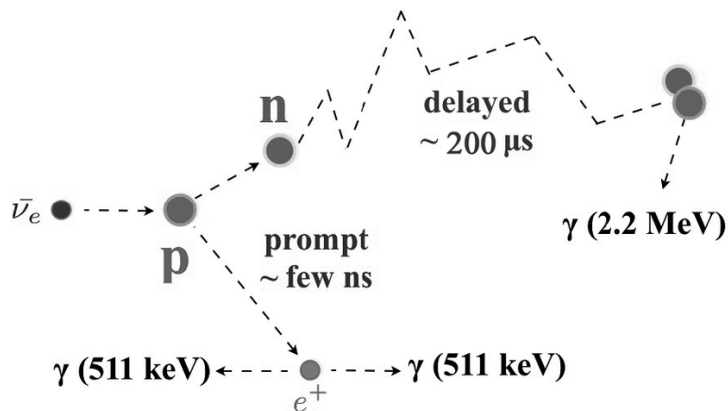
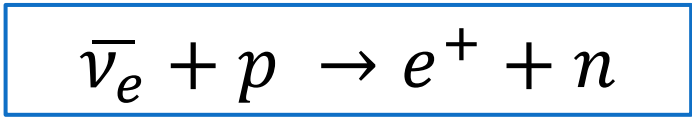
# Reactor neutrino detection



In nuclear reactors,  $\bar{\nu}_e$  are emitted from  $\beta$ -decays of fission fragments:

- ▶ typical flux  $\sim 2 \times 10^{20} \bar{\nu}_e / \text{s} / \text{GW}_{\text{th}}$
- ▶ most reactor  $\bar{\nu}_e$  have energy  $E < 10 \text{ MeV}$
- ▶  $> 99\%$  of reactor  $\bar{\nu}_e$  emissions come from fissions of four main isotopes:  $^{235}\text{U}$ ,  $^{239}\text{Pu}$ ,  $^{241}\text{Pu}$ ,  $^{238}\text{U}$ .

$\bar{\nu}_e$  can be detected via the Inverse Beta Decay (IBD) reaction:



- ▶ Energy threshold: 1.806 MeV
- ▶ Prompt visible energy:  
 $E_{\text{vis}} = E_{\nu} - 0.784 \text{ MeV}$
- ▶ Time and space coincidence between prompt and delayed signal allows to reject uncorrelated background

# JUNO: a multipurpose experiment

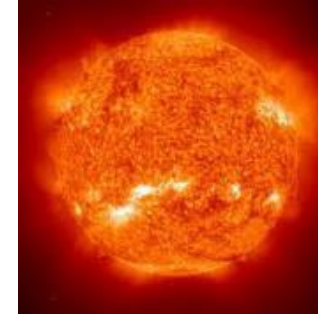


- ▶ The JUNO detector will measure neutrinos from different sources in the energy range from  $\sim$ MeV to tens of GeV
- ▶ JUNO will offer exciting opportunities for addressing many important topics in neutrino and astro-particle physics

JUNO collaboration, [Prog.Part.Nucl.Phys. 123 \(2022\) 103927](#)



Reactor  $\nu$   
 $\sim$  50 IBD/day

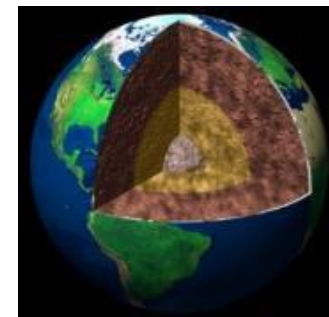
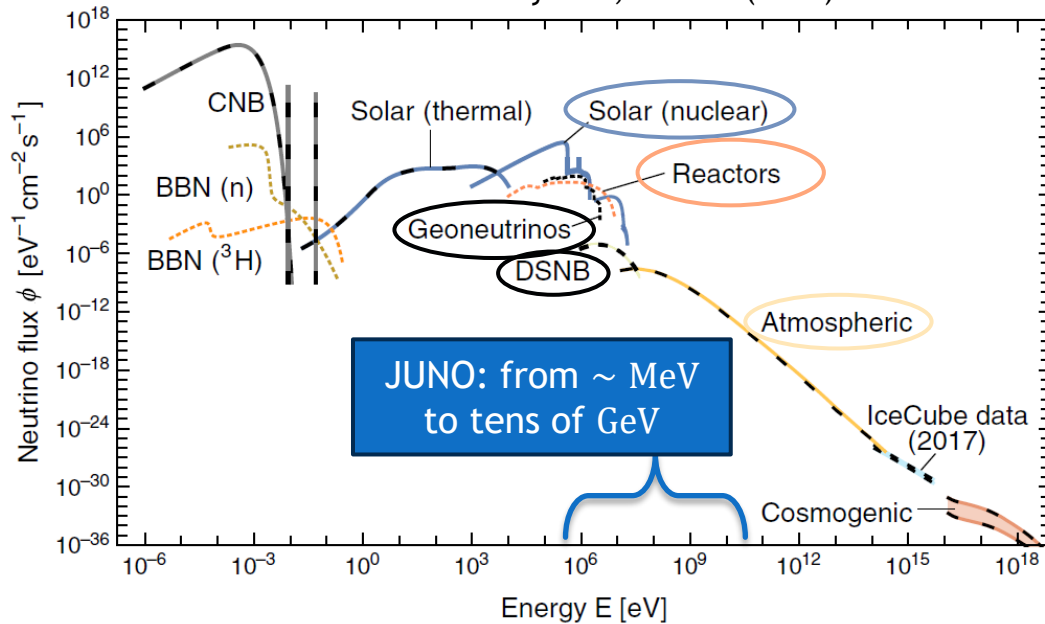


Solar  $\nu$   
 hundreds/day

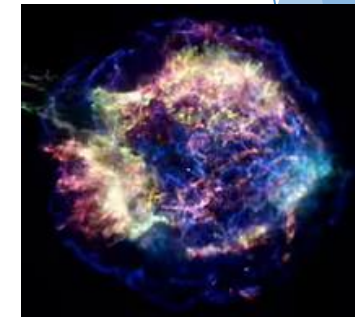


Atmospheric  $\nu$   
 several/day

Grand unified neutrino spectrum at Earth  
 Rev. Mod. Phys. 92, 045006 (2020)



Geoneutrinos  
 $\sim$  400 IBD/yr

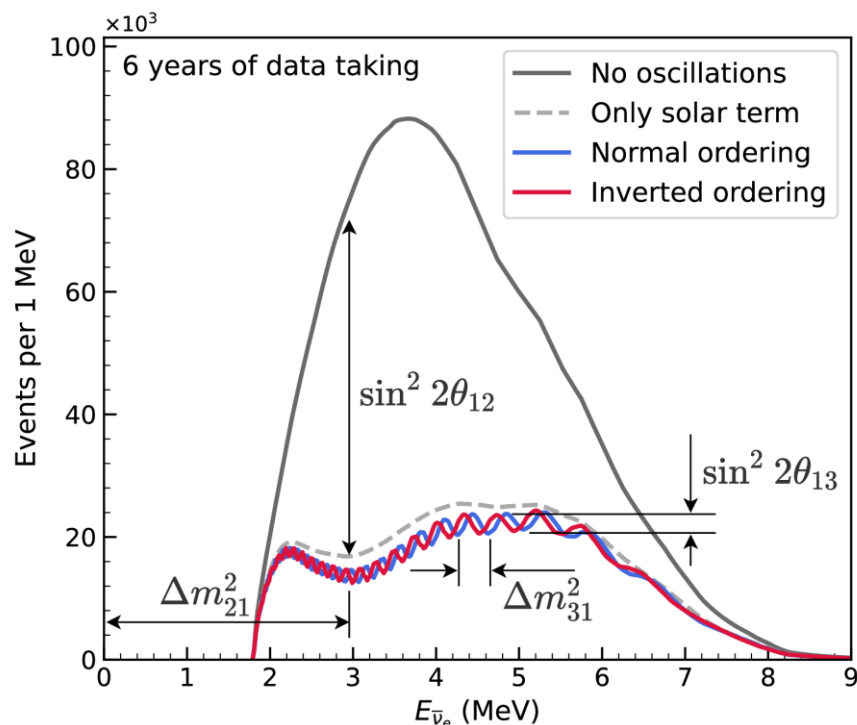


Supernova  $\nu$   
 $\sim$  5000 IBD in 10 s for CCSN at 10 kpc

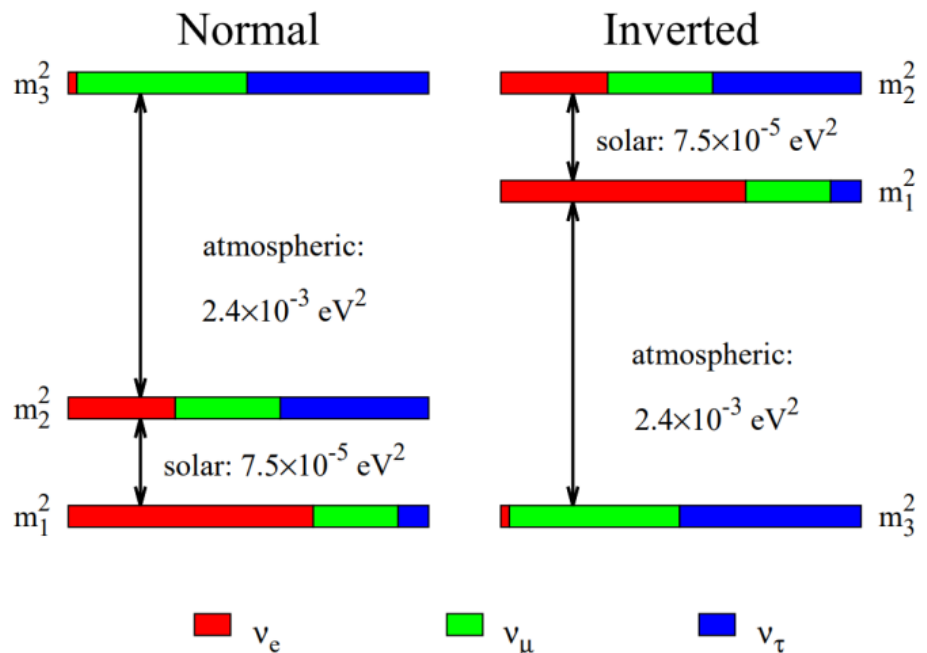
# The JUNO primary physics goals



- ▶ Determine the Neutrino Mass Ordering (NMO)
- ▶ Measure three oscillation parameters ( $\sin^2 \theta_{12}$ ,  $\Delta m_{21}^2$ ,  $\Delta m_{31}^2$ ) with sub-percent precision



JUNO collaboration, [CPC 46, 123001 \(2022\)](#)



JUNO will be the first experiment to observe both fast ( $\sin^2 \theta_{13}$ ,  $\Delta m_{31}^2$ ) and slow ( $\sin^2 \theta_{12}$ ,  $\Delta m_{21}^2$ ) oscillations in vacuum



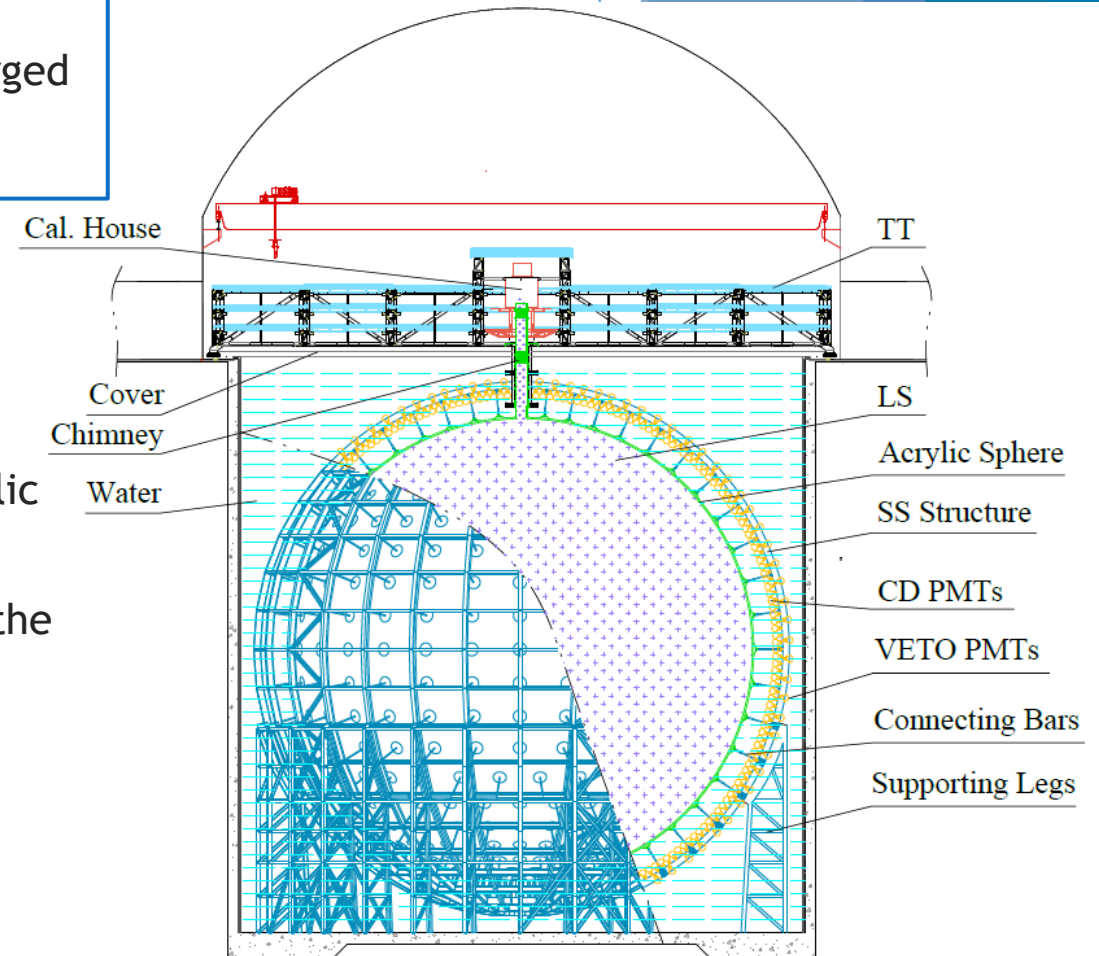
The interference pattern depends on NMO

# The JUNO detector



- ❑ 20 kton liquid scintillator (LS) detector
- ❑ Water Cherenkov detector in which the LS detector is submerged
- ❑ Plastic scintillator array on top (Top Tracker)

- ▶ The **20 kton LS** is contained in an **Acrylic Sphere**
  - Inner Diameter (ID): 35.4 m
  - Thickness: 12 cm
- ▶ A Stainless Steel (**SS**) **Structure** (40.1 m ID) supports the acrylic sphere via 590 connecting bars
- ▶ The light emitted by the LS is detected by **PMTs** installed on the inner surface of the SS structure
  - 17612 20-inch PMTs (75.2% photocathode coverage)
  - 25600 3-inch PMTs (2.7% photocathode coverage)
- ▶ The **Water pool** is filled with 35 kton of ultrapure water
  - ID: 43.5 m, Height: 44 m
  - 2400 20-inch PMTs installed on the outer surface of the SS structure



# The JUNO detector construction



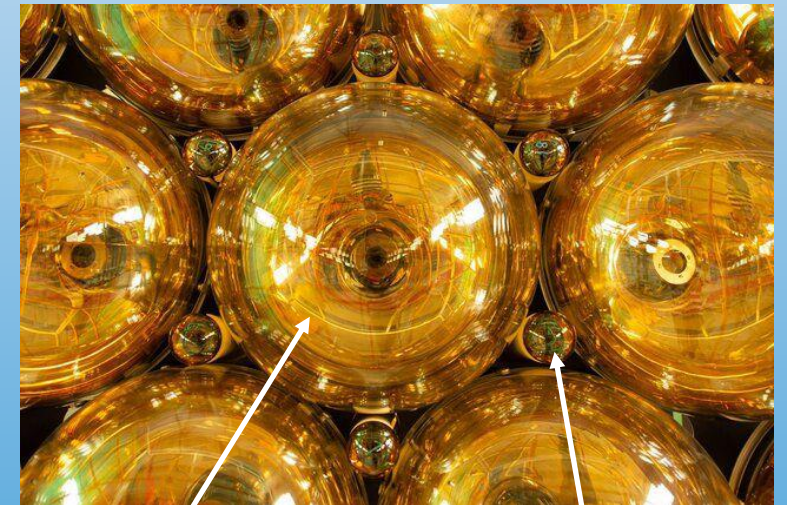
- ▶ SS structure built from bottom to top



- ▶ Then, acrylic sphere built from top to bottom, layer by layer



Production and testing done for all PMTs, installation close to completion



20-inch PMTs  
(large)

3-inch PMTs  
(small)

# The JUNO key challenges



## 1. Large statistics

- huge scintillator mass ( $\sim 1.5 \times 10^{33}$  target protons)
- nuclear reactor power: 26.6 GW<sub>th</sub> (6 reactors 2.9 GW<sub>th</sub> each at Yangjiang + 2 reactors 4.6 GW<sub>th</sub> each at Taishan)

## 2. Energy resolution: 2.95% @ 1MeV

[arXiv:2405.17860](https://arxiv.org/abs/2405.17860)

Uncertainty on the intrinsically non-linear energy scale:  $\leq 1\%$

## 3. Low background:

- underground laboratory
- scintillator purification system
- material screening to meet the requirements for Th/U/K contaminations:
  - $< 10^{-15}$  g/g in LS and  $< 10^{-12}$  g/g in acrylic
- veto systems (Top Tracker, Water Cherenkov detector )

## 4. Knowledge of unoscillated reactor spectra



Taishan Antineutrino Observatory (TAO)



# Expected $\bar{\nu}_e$ spectrum at JUNO



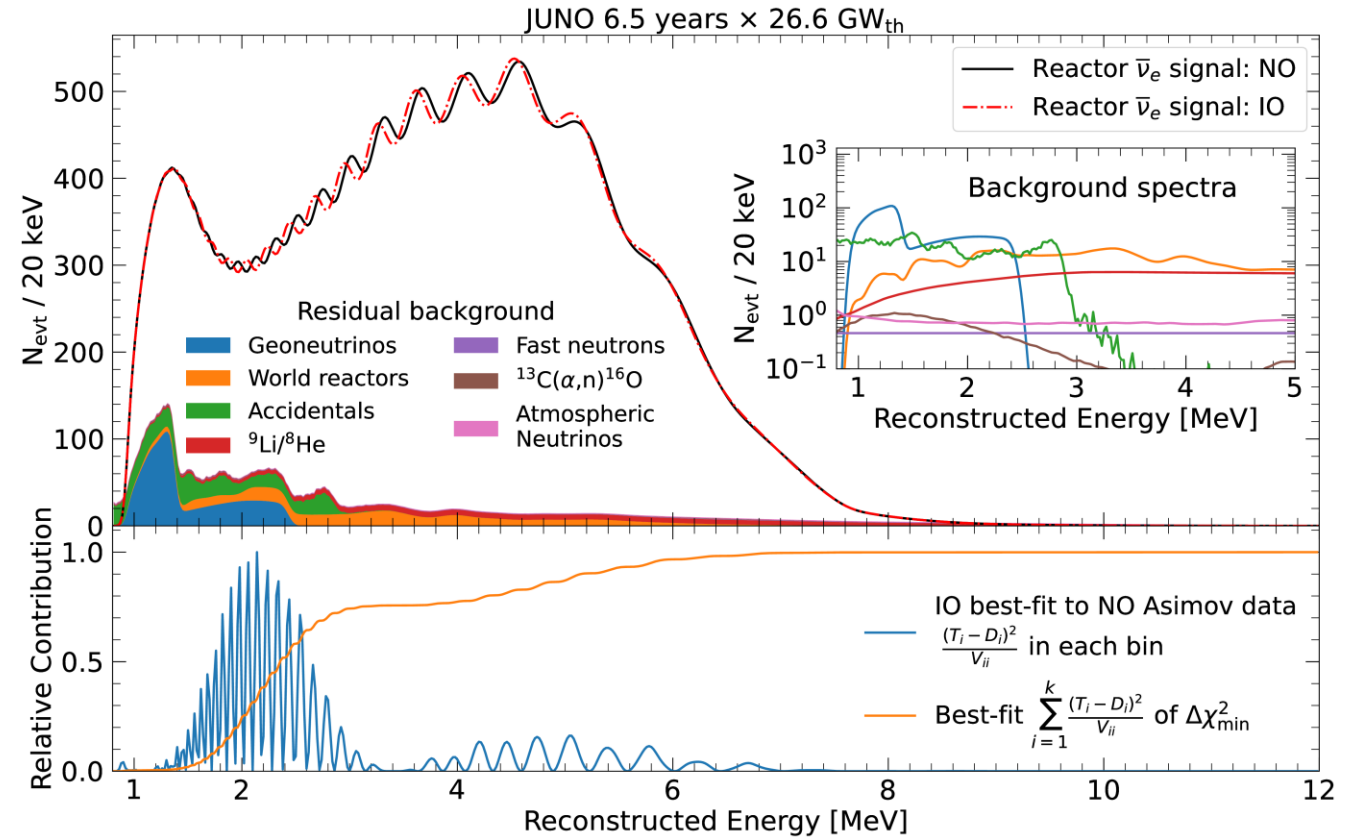
Top panel:

- ▶ Energy spectra in both the NO and IO scenarios without any statistical or systematic fluctuations
- ▶ The background spectra in the main figure are stacked on top of each other

Bottom panel:

- ▶ Relative contribution to the  $\Delta\chi^2$  obtained when fitting the IO spectrum with the NO hypothesis.

The most sensitive region for JUNO's NMO determination is in the visible energy range from 1.5 to 3 MeV



A  $3\sigma$  median sensitivity to reject the wrong mass ordering hypothesis can be reached with an exposure of about 6.5 years  $\times$  26.6 GW thermal power

[arxiv:2405.18008](https://arxiv.org/abs/2405.18008)

# The TAO experiment

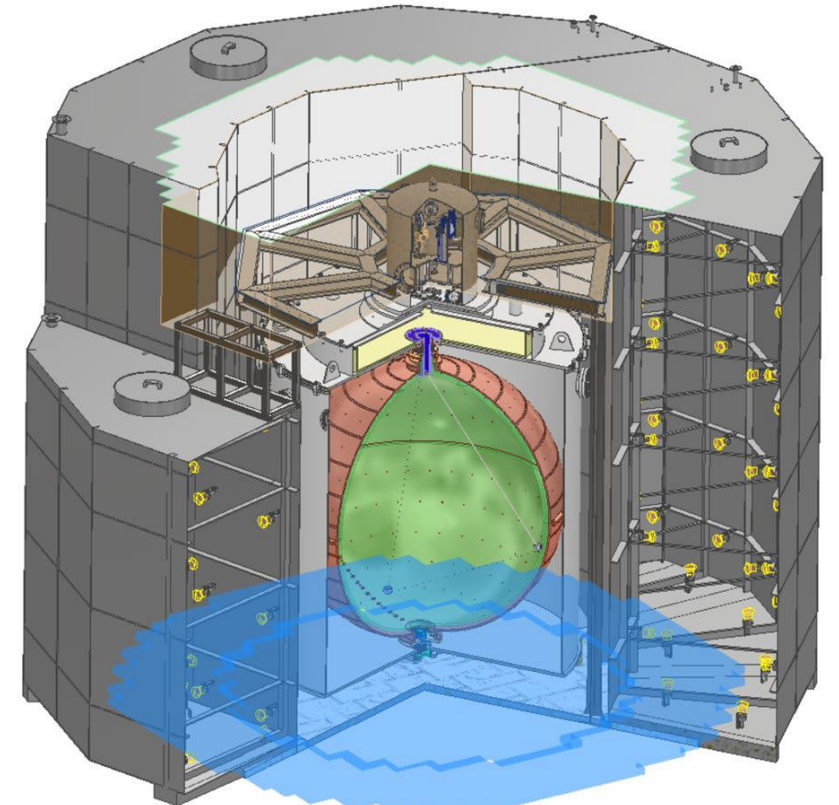


TAO (Taishan Antineutrino Observatory) is a satellite experiment of JUNO consisting of a ton-level liquid scintillator detector:

- ▶ positioned at **44 m** from the core of the Taishan-1 reactor and at 217 m from Taishan-2;
- ▶ will measure the  $\bar{\nu}_e$  spectrum with unprecedented **energy resolution** (better than **2% at 1 MeV**);
- ▶ will observe **~1000 IBDs/day** in the fiducial volume.

The main goals and physics potential of TAO are:

- ▶ provide a **reference spectrum** for JUNO;
- ▶ **benchmark for** future reactor neutrino **experiments** and **nuclear databases** (fine structures observation);
- ▶ measurement of isotopic IBD yields with larger sampled range of fission fractions;
- ▶ light sterile neutrino searches;
- ▶ increase the reliability and verify the technology for reactor monitoring.



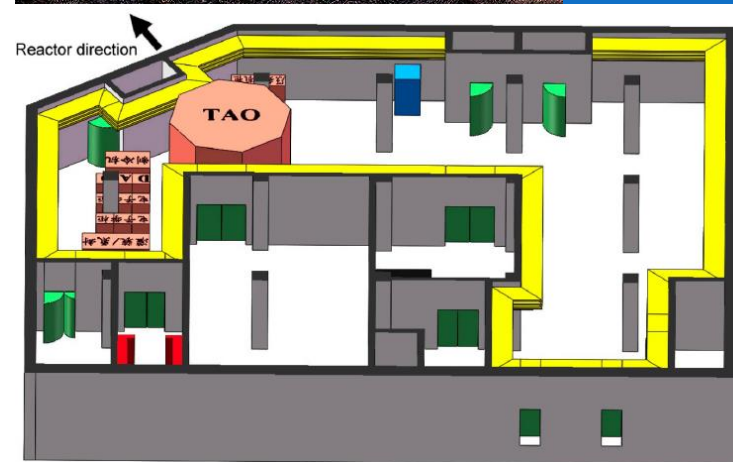
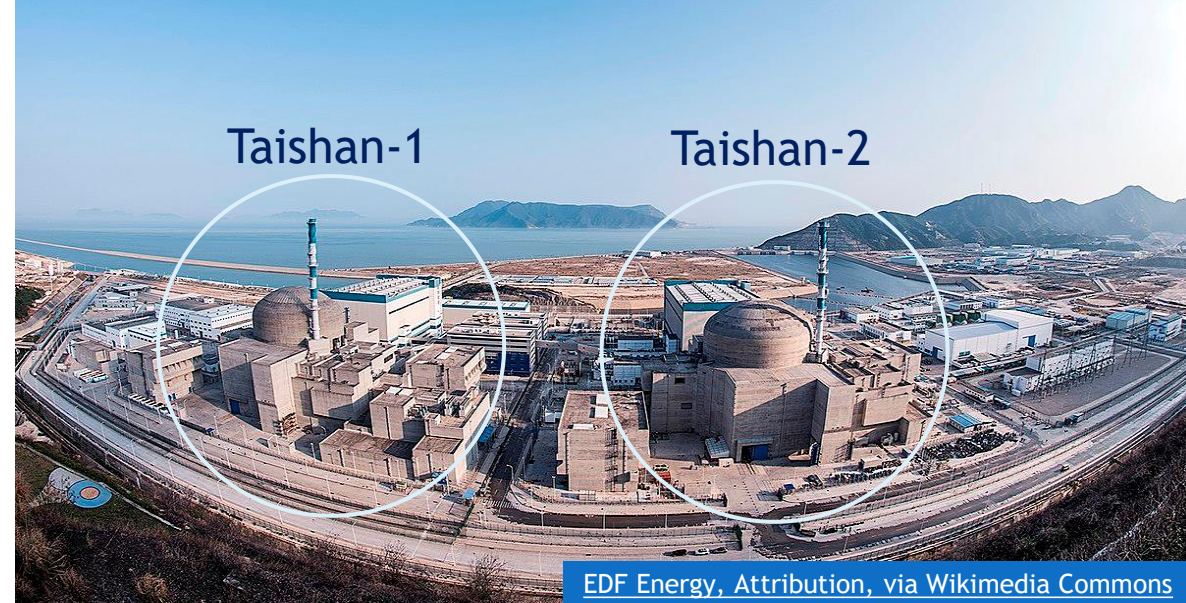
# TAO at the Taishan NPP



- ▶ Taishan Nuclear Power Plant has two cores currently in operation (other two cores might be built later)
- ▶ Both reactors are European Pressurised Reactor (EPR) with  $4.6 \text{ GW}_{\text{th}}$  thermal power
- ▶ Taishan-1 reached first criticality and was connected to the grid in June 2018  
→ the first running EPR in the world!

- ▶ The TAO detector will be installed in a basement at 9.6 m underground, outside of the concrete containment shell of the reactor core
  - $>99.99\%$  signal from Taishan-1+Taishan-2
  - 4% signal from Taishan-2
- ▶ Muon rate and cosmogenic neutron rate are measured to be 1/3 of those on the ground

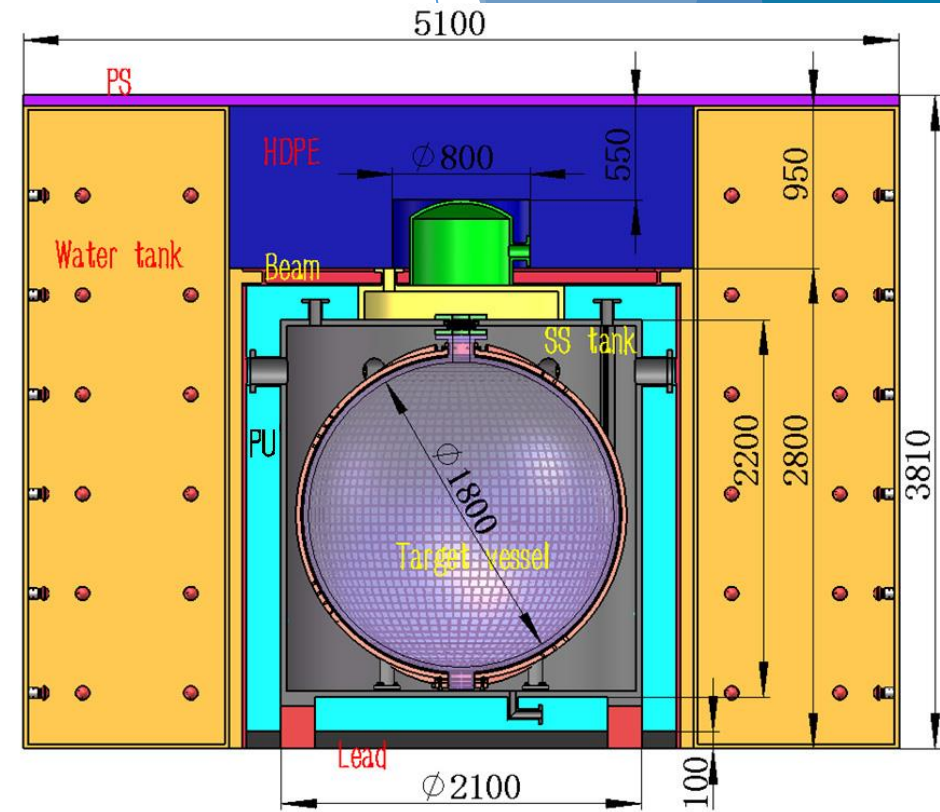
Taishan Nuclear Power Plant (NPP), Guangdong, China



# The TAO detector



- ▶ 2.8 ton Gadolinium-doped Liquid Scintillator (**Gd-LS**) filled in a spherical **acrylic vessel** of 1.8 m in inner diameter (*ID*).
- ▶ Scintillation and Cherenkov light is detected by **4024** Silicon Photomultipliers (**SiPM**)  $5.08\text{ cm} \times 5.08\text{ cm}$  each, covering  $\sim 10\text{ m}^2$  with 50% photon detection efficiency.
- ▶ SiPM tiles on the inner surface of a spherical **copper shell** (*ID* = 1.882 m)
- ▶ Cylindric **stainless steel tank** ( $OD = 2.1\text{ m}$ ,  $H = 2.2\text{ m}$ ) filled with Linear Alkylbenzene (LAB) as buffer liquid to shield the radioactivity, stabilize the temperature, and optically couple acrylic and SiPMs.
- ▶ SS tank is insulated with 20 cm thick Polyurethane (PU) to operate at  $-50^\circ\text{C}$   $\rightarrow$  reduce the dark noise of SiPMs to  $\sim 100\text{ Hz/mm}^2$ .
- ▶ The central detector is surrounded by:
  - 1.2 m thick water tanks on the sides
  - 1 m High Density Polyethylene (HDPE) on the top
  - 10 cm lead at the bottomto shield the ambient radioactivity and cosmogenic neutrons.



- ▶ Muon veto system: water tank instrumented with PMTs + plastic scintillator array on top

# IBD detection and event selection in TAO



Inverse beta decay (IBD) in LS with 0.1% Gd-loading:

- ▶ 87% neutron captures by Gd (and 13% by hydrogen)
- ▶ Average capture time by Gd is about 30  $\mu$ s

Event selection:

- ▶ Fiducial volume cut: 25 cm from the LS boundary to reduce energy leakage and mitigate backgrounds  $\rightarrow$  1 ton fiducial mass
- ▶ Delayed energy cut:  $7 \text{ MeV} < E_d < 9 \text{ MeV}$  to reduce background rate by one order of magnitude
- ▶ Prompt energy cut:  $E_p > 0.9 \text{ MeV}$
- ▶ Prompt-delayed time coincidence cut:  $1 \mu\text{s} < \Delta t < 100 \mu\text{s}$

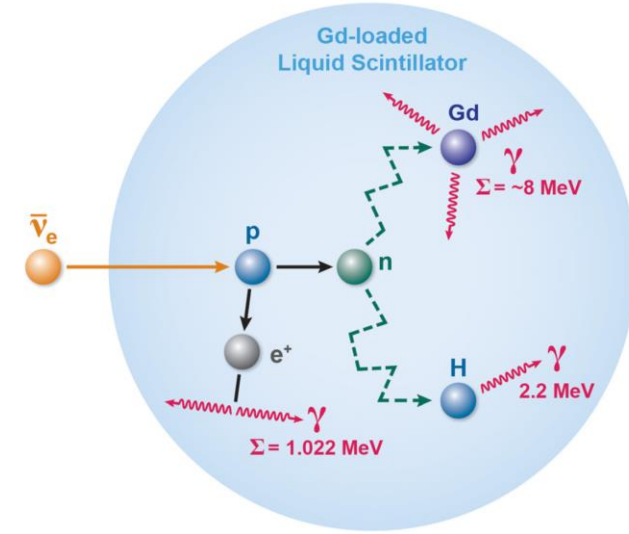
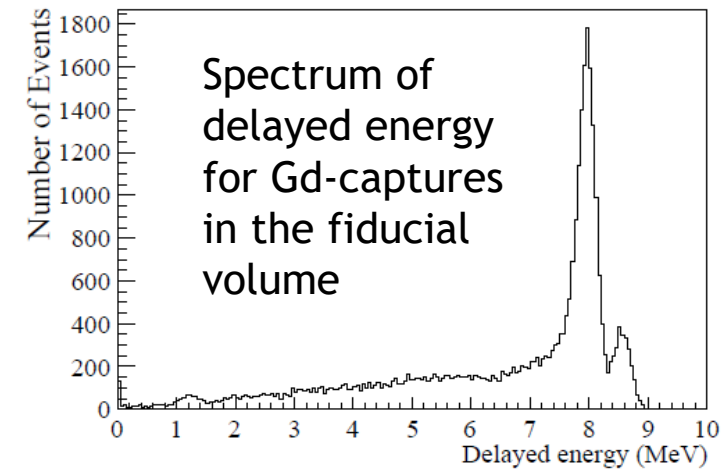


Image Credit:  
Rep. Prog. Phys.  
82, 036201 (2019)

	Efficiency
Captures by Gd	87%
Delayed energy cut	59%
Prompt energy cut	99.8%
Time coincidence cut	97%

Overall detection efficiency  $\sim 50\%$

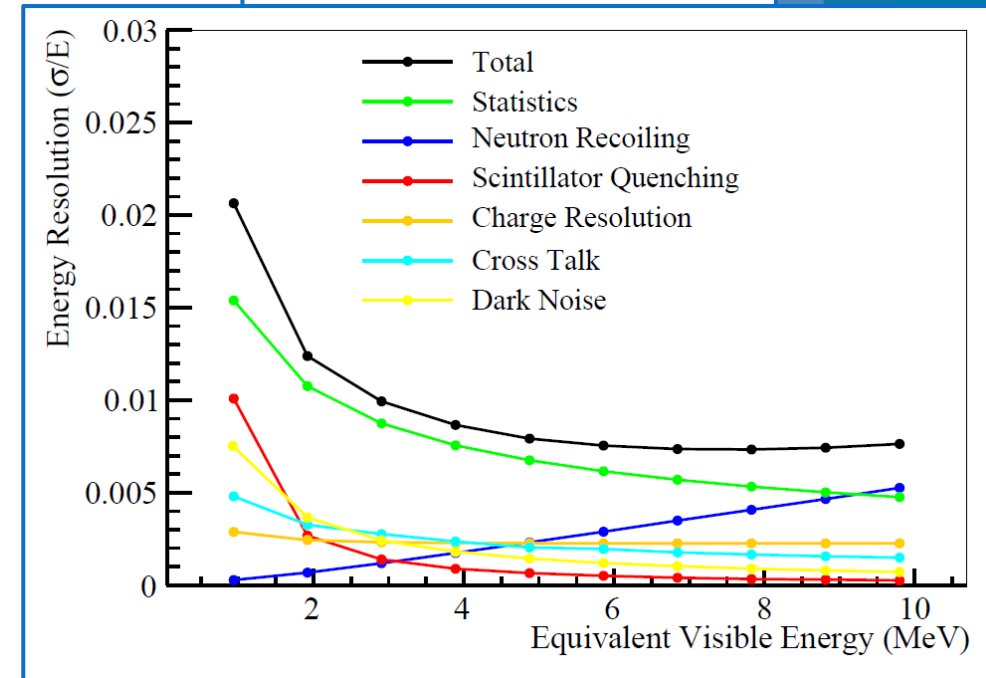


# TAO energy resolution



	JUNO	TAO
Coverage	~ 75%	~ 94%
Photon detection efficiency	~ 30%	> 50%
Attenuation length	> 20 m ( $R = 17.2$ m)	> 20 m ( $R = 0.9$ m)
Photoelectron yield	~ 1665 PE/MeV	~ 4500 PE/MeV
Energy resolution	2.95% @ 1 MeV	~ 2% @ 1 MeV

TAO energy resolution



Non-stochastic effects affecting energy resolution in TAO:

- ▶ at low energies, the contribution from the LS quenching effect might be quite large;
- ▶ at high energies, the smearing from neutron recoil of IBD becomes dominant.

In most of the energy region of interest, the energy resolution of TAO will be sub-percent!

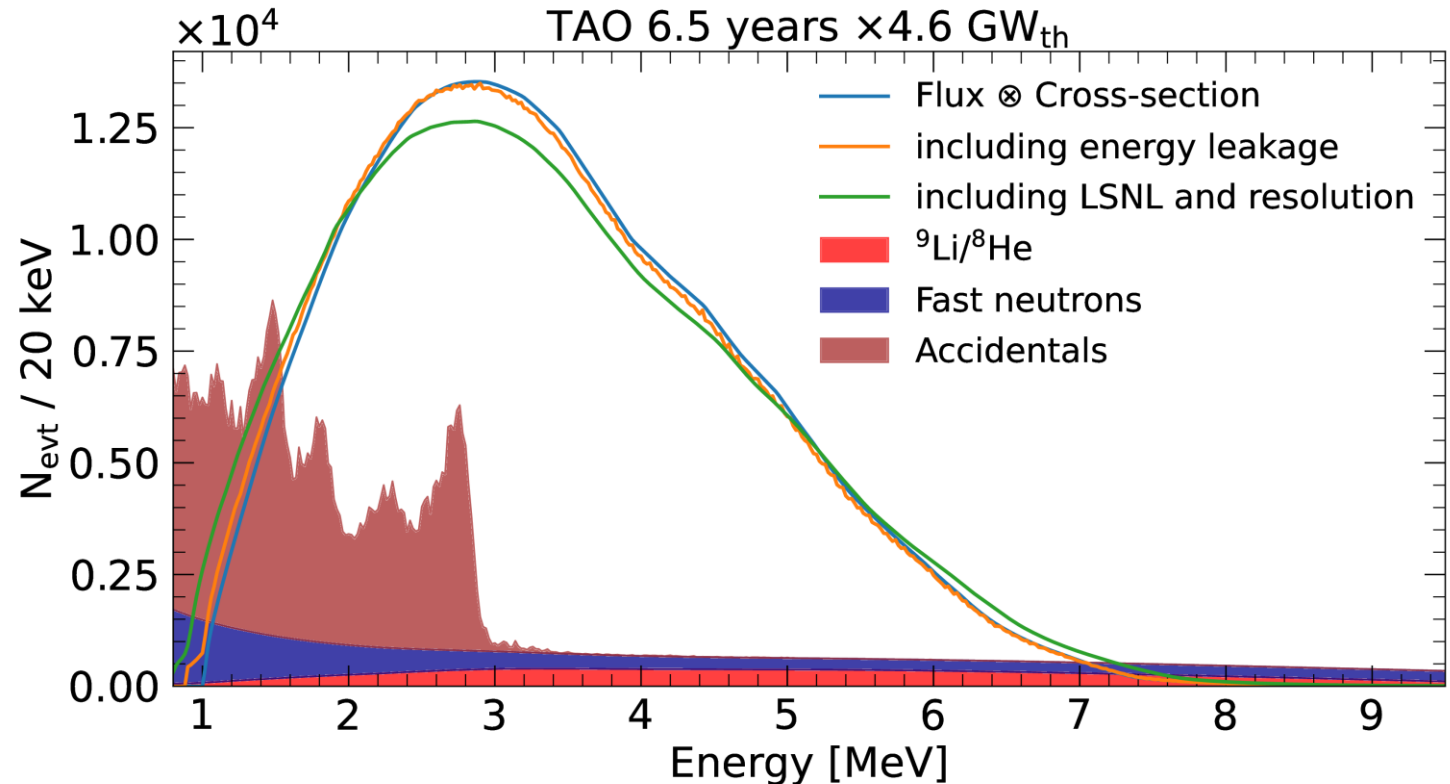
# TAO expected signal and background

▶ Signal spectrum: 
$$S(E_\nu) = \frac{N_p \epsilon \sigma(E_\nu)}{4\pi L^2} \phi(E_\nu)$$

▶ The signal spectrum is shown w/ and w/o applying energy leakage, liquid scintillator non-linearity (LSNL), and energy resolution effects.

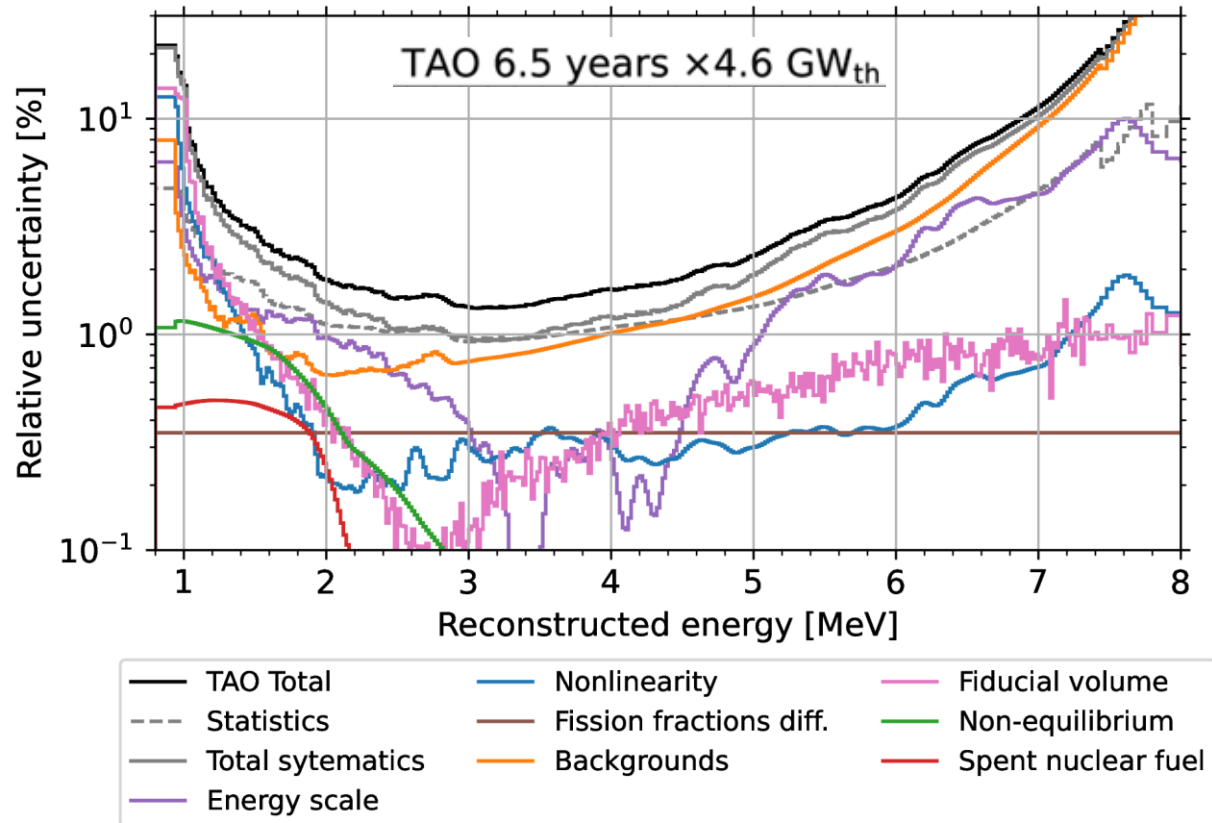
▶ TAO backgrounds will be directly measured exploiting the reactor-off data (about one month per year)

Type	Rate [day <sup>-1</sup> ]
Signal	1000
Fast neutron	86
<sup>9</sup> Li/ <sup>8</sup> He	54
Accidental	190



# Shape uncertainty of TAO spectrum

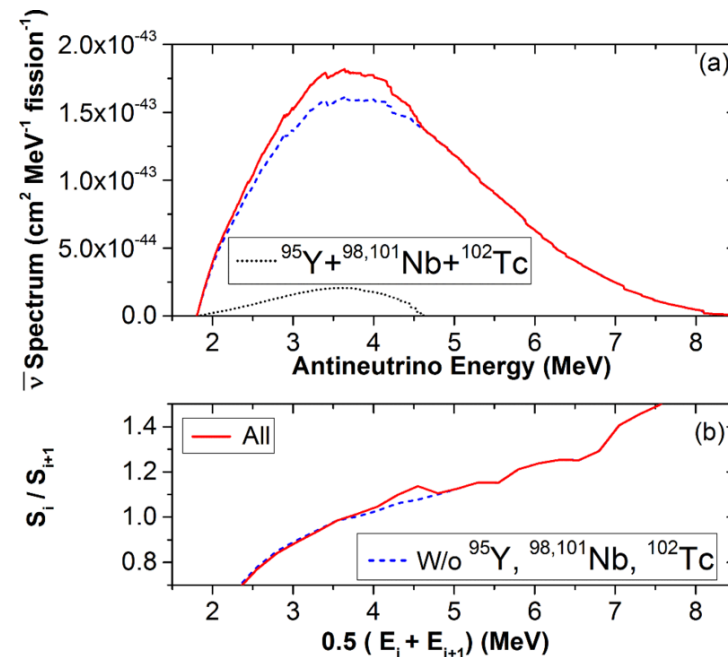
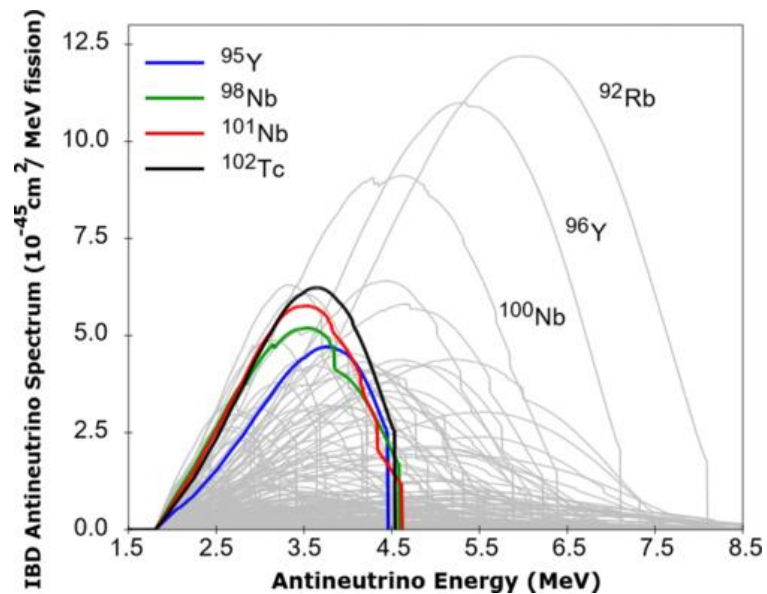
- ▶ Statistical uncertainty:  $\sim 1\%$  in the energy range 2 – 5 MeV (20 keV bin width)
- ▶ Systematic uncertainties are not negligible at low/high energies, but are  $\sim 1\%$  in the central energy range 2 – 5 MeV





# Fine structure measurement

- ▶ The reactor  $\bar{\nu}_e$  spectrum is composed of spectra from thousands of beta decay branches.
- ▶ The end point of each  $\bar{\nu}_e$  spectrum has a sharp edge (Coulomb correction), which produces a percent-level fine structure.
- ▶ Thanks to its excellent energy resolution, TAO will uncover the fine structures of reactor  $\bar{\nu}_e$  spectrum for the first time.
- ▶ The TAO measurement will provide a benchmark to test nuclear databases, comparing the experimental data with the predictions of the summation method.



Plot taken from A. A. Sonzogni, M. Nino, and E. A. McCutchan, [Phys. Rev. C 98, 014323 \(2018\)](#)

# Conclusions



- ▶ JUNO detector is in an advanced construction status and is expected to start data taking next year.
- ▶ JUNO has a rich physics program in particle and astroparticle physics.
- ▶ Main goal: identification of NMO and precise measurement of oscillation parameters thanks to its unprecedented energy resolution.
- ▶ TAO 1:1 prototype has been successfully tested at IHEP and, after disassembling and re-installation in Taishan, will start data taking next year.
- ▶ TAO will provide a high-resolution reference reactor spectrum for JUNO NMO analysis, but also precious experimental data to be used as benchmark for the modelling of reactor antineutrino spectrum.