

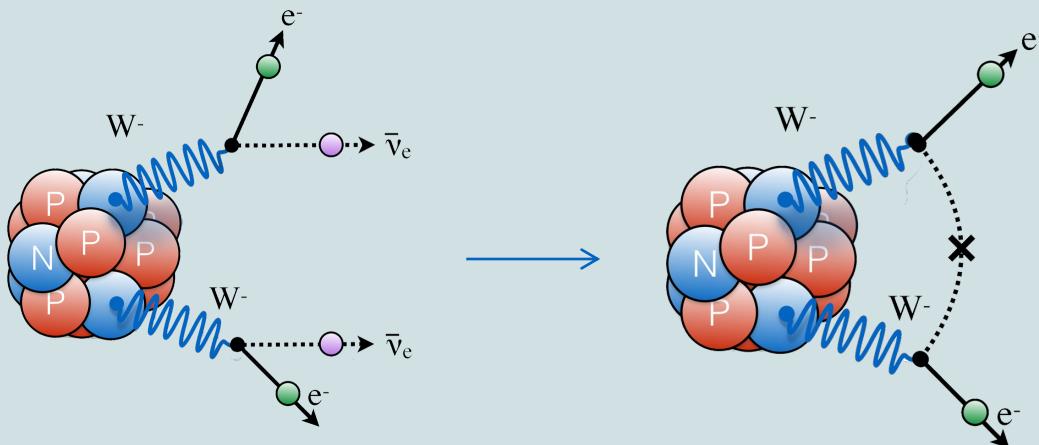
# Latest results from KamLAND-Zen



The search for neutrinoless double beta decay ( $0\nu\beta\beta$ ) with the complete dataset

# Observation of $0\nu\beta\beta$ to confirm neutrinos' majorana nature

$$2\nu\beta\beta : (A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\cancel{\bar{\nu}_e} \quad 0\nu\beta\beta \quad (T_{1/2}^{0\nu})^{-1} \propto \langle m_{\beta\beta} \rangle^2$$

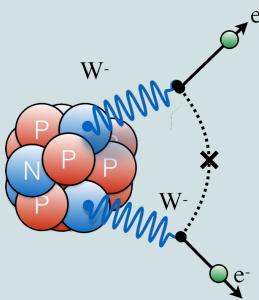
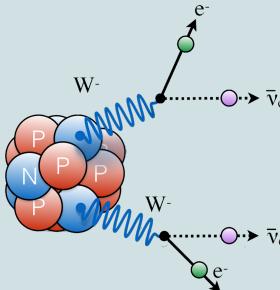


**Majorana ( $\nu = \bar{\nu}$ ) relevant for**

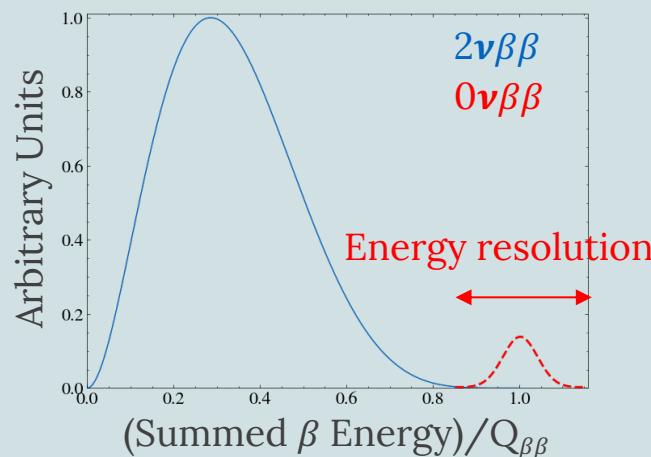
- Neutrino mass via the seesaw mechanism
- Matter-antimatter asymmetry via leptogenesis

$2\nu\beta\beta$  exceptionally slow nuclear process,  $T_{1/2} \sim 10^{18-24}$  years

# Expected signal for neutrinoless double beta decay



Energy distribution of created electrons



Abundance: 8.86%

Q-value: 2.458 MeV

Most stringent limit on  $0\nu\beta\beta$  in **xenon** from KamLAND-Zen:  $T_{1/2} > 3.8 \times 10^{26}$  years

# KamLAND Collaboration

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# General purpose detector

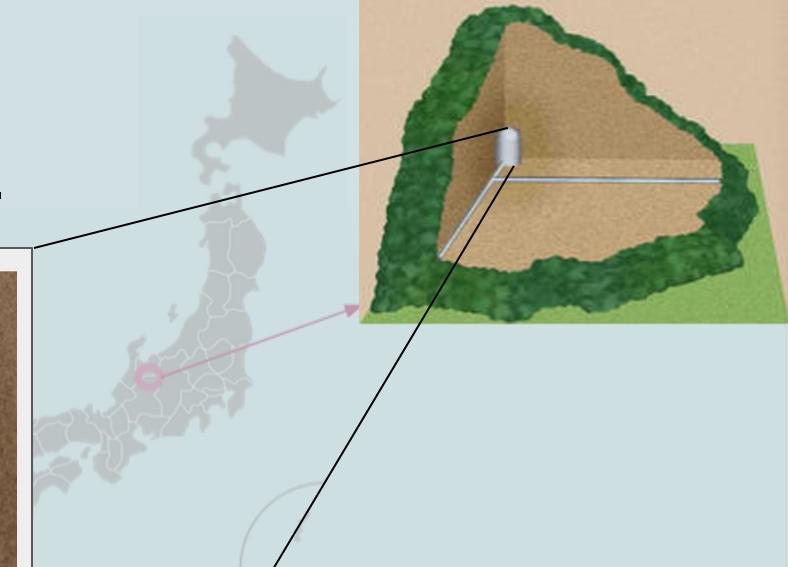
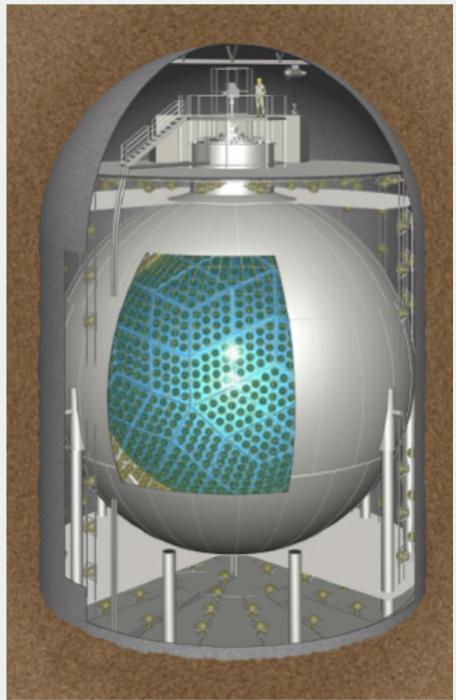
Solar neutrinos

Geo and reactor  
neutrinos

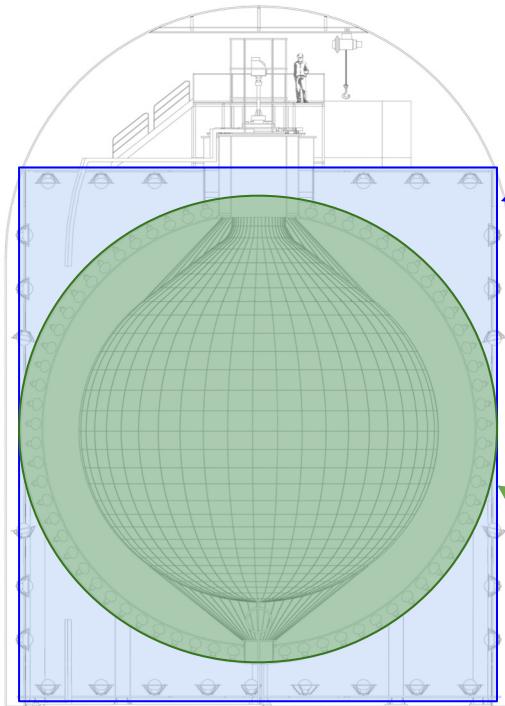
Accelerator neutrinos

Astrophysical  
neutrinos

**Neutrinoless double  
beta decay**



# Kamioka Liquid Scintillator Antineutrino Detector



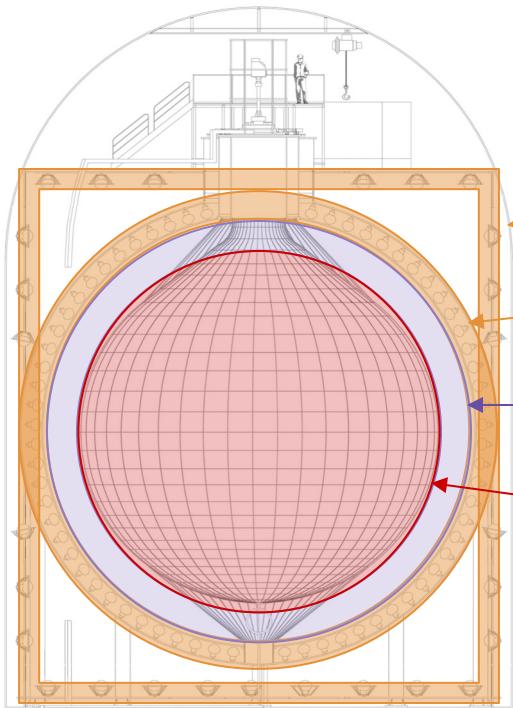
Outer Cherenkov detector

Inner Scintillation detector

Cylindrical tank,  $\varnothing 18\text{m}$

- 3.2kt pure water

# Kamioka Liquid Scintillator Antineutrino Detector



Outer Cherenkov detector

Outer PMTs

Inner PMTs

Buffer oil

Liquid Scintillator balloon

Inner Scintillation detector

Cylindrical tank,  $\varnothing 18\text{m}$

- 3.2kt pure water

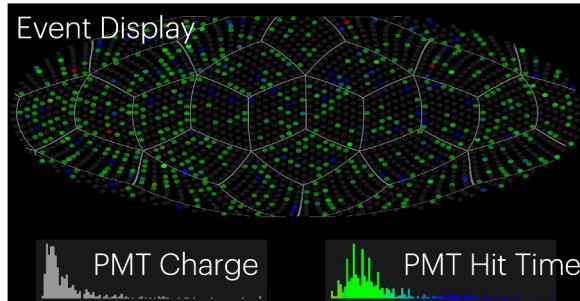
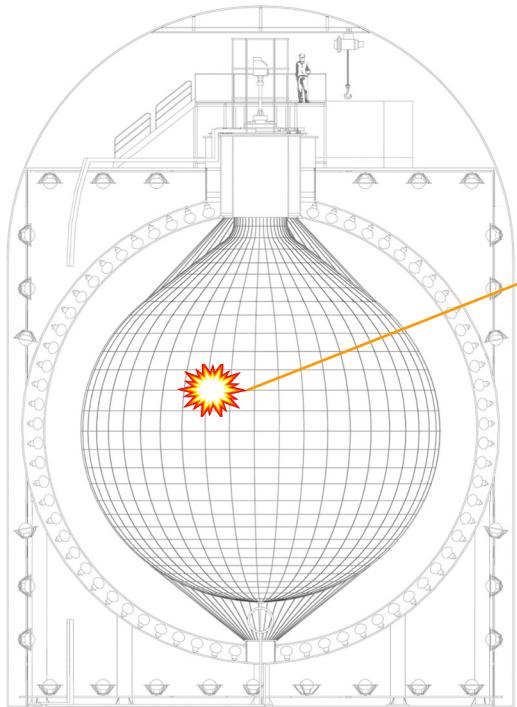
$\sim 1800$  17- & 20-inch PMTs

Non-scintillation oil 1.4kt

**KamLAND-LS,  $\varnothing 13\text{m}$ , 1kt**

- Dodecane 80.2%
- Pseudocumene 19.8%
- PPO 1.36g/L

# Kamioka Liquid Scintillator Antineutrino Detector



Charge & Timing

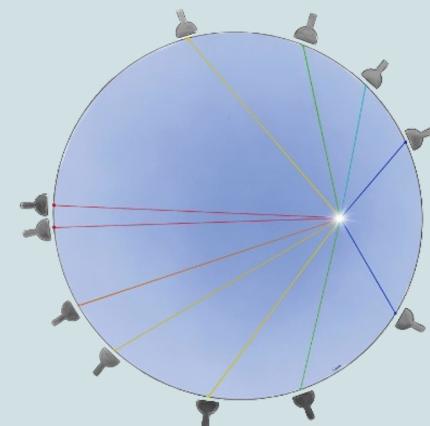


Energy & Vertex

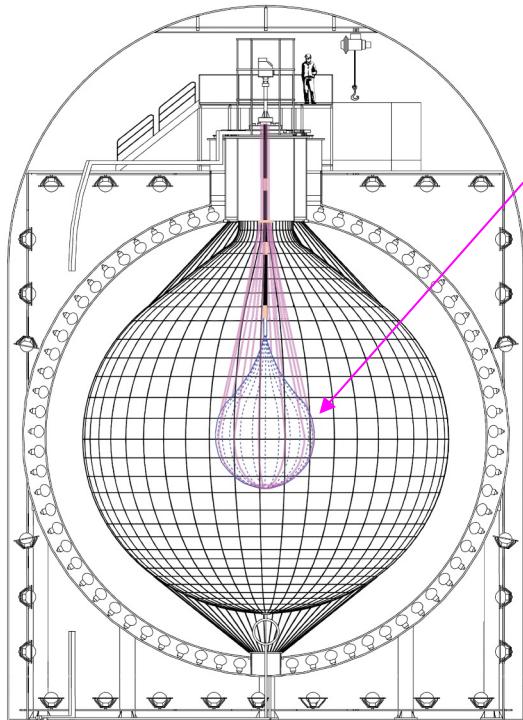
$$\frac{6.7\%}{\sqrt{E(\text{MeV})}} \quad \frac{13.7\text{cm}}{\sqrt{E(\text{MeV})}}$$

## Inner detector

- ~ 1300 17-inch PMTs
- ~ 550 20-inch PMTs



# KamLAND-Zen: Zero Neutrino Double Beta



## Inner Balloon Xe-LS

Zen 400 (Oct. 2011 – Oct. 2015)

Phase I: PRL 110 (2013): 0625023

Phase II: PRL 117.8 (2016): 082503

Zen 800 (Feb. 2019 – Jan. 2024)

First dataset: PRL 130.5 (2023): 051801

Complete: arXiv preprint 2406.11438

First → Complete dataset

Livetime 523 → **1131** days

Exposure 0.97 → **2.1** ton yr  $^{136}\text{Xe}$



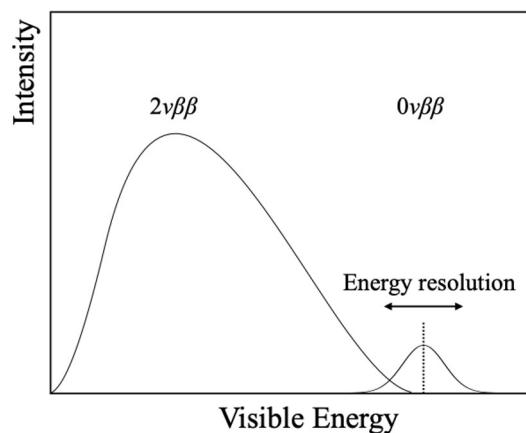
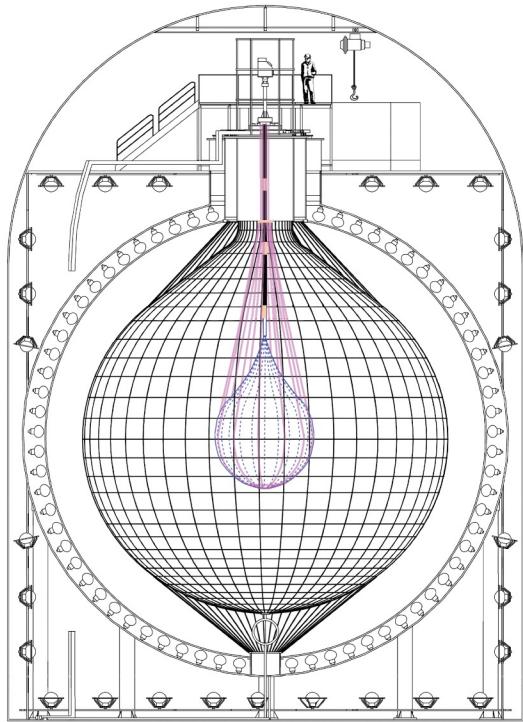
Xe-LS balloon,  $\varnothing 3.8\text{m}$ , 24t

- 3.13% enriched xenon
- $\rightarrow$  **745 kg Xe**
- 90.85%  $^{136}\text{Xe}$

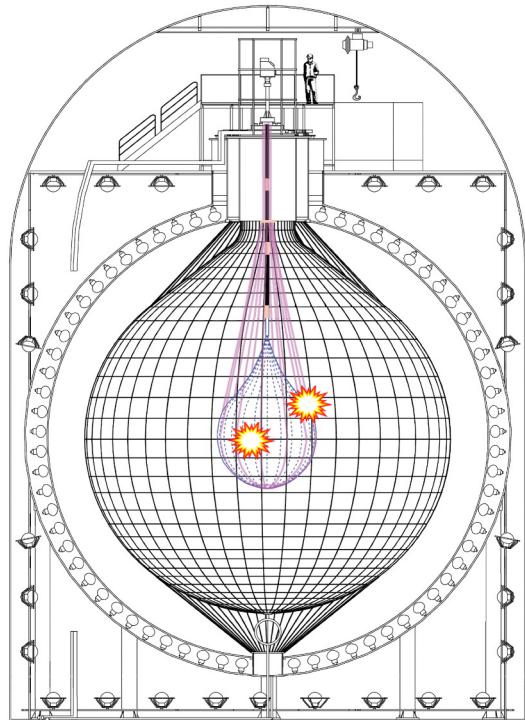
$$Q_{\beta\beta} = 2.458 \text{ MeV}$$

# Dominant backgrounds for the $0\nu\beta\beta$ search

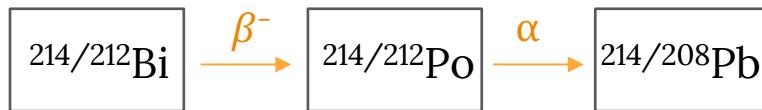
➤  $2\nu\beta\beta$  decay



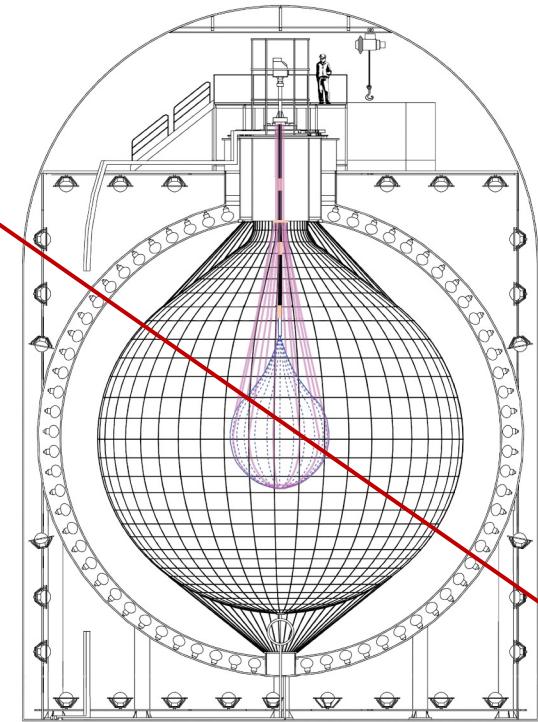
# Dominant backgrounds for the $0\nu\beta\beta$ search



- $2\nu\beta\beta$  decay
- Radioactive Impurities (RI)
  - In Xe-LS:  $^{232}\text{Th}$  ( $97.9 \pm 0.5\%$ ) &  $^{238}\text{U}$  ( $99.95 \pm 0.03\%$ ) tagging efficiency
  - External
    - Dominated by  $^{238}\text{U}$  because delayed coincidence  $^{214}\text{Bi}(\beta^-)$  -  $^{214}\text{Po}(\alpha)$  efficiency  $\sim 50\%$  due to absorption of  $\alpha$  on non-scintillating IB

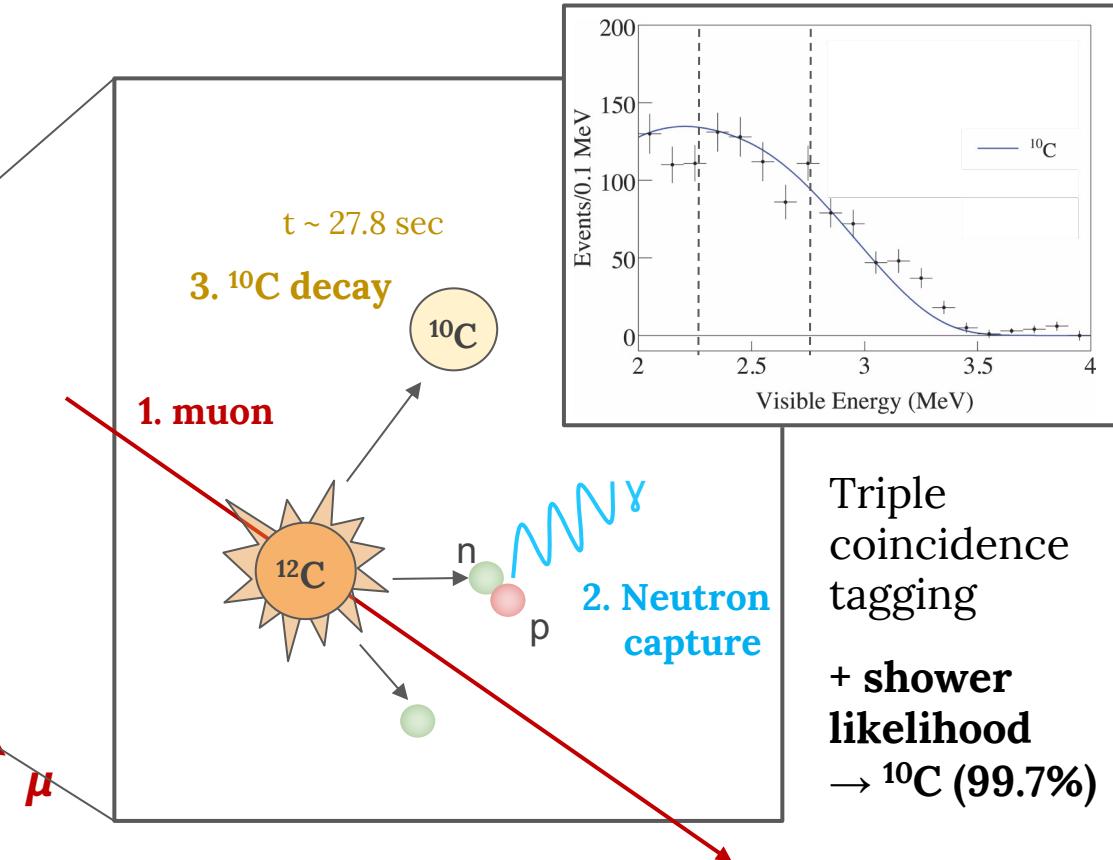
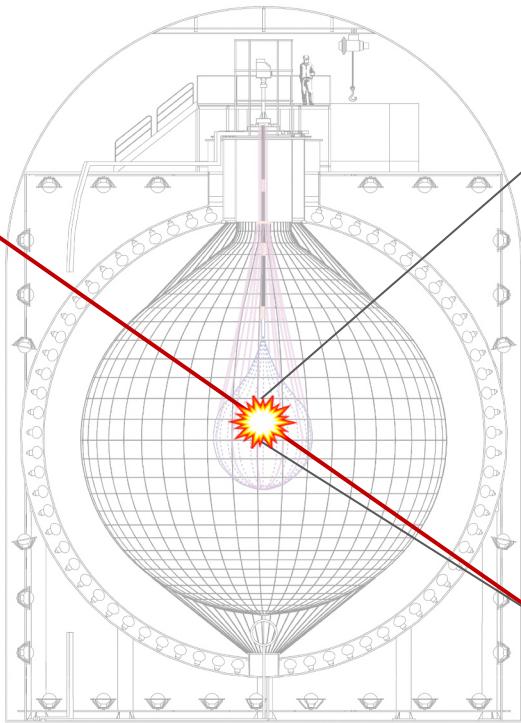


# Dominant backgrounds for the $0\nu\beta\beta$ search



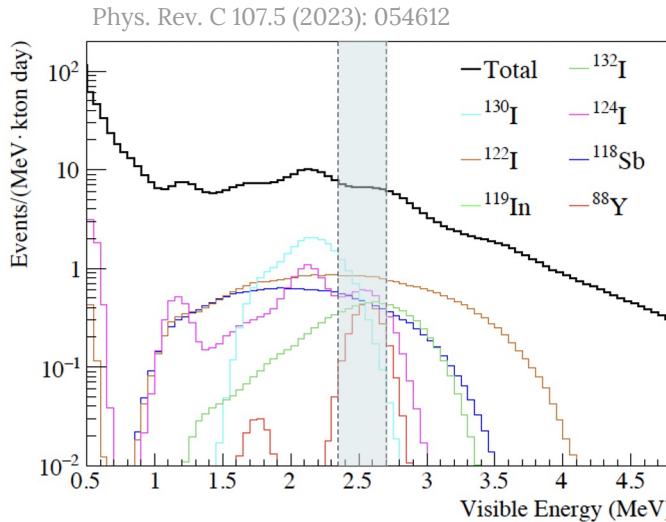
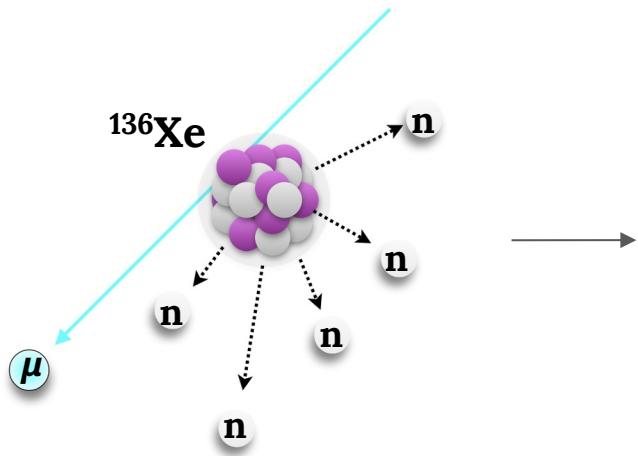
- $2\nu\beta\beta$  decay
- Radioactive Impurities (RI)
  - In Xe-LS:  $^{232}\text{Th}$  ( $97.9 \pm 0.5\%$ ) &  $^{238}\text{U}$  ( $99.95 \pm 0.03\%$ ) tagging efficiency
  - **External:** IB material
- Muon spallation products
  - Short-lived: carbon spallation products dominated by  $^6\text{He}$ ,  $^8\text{Li}$ ,  $^{10}\text{C}$ ,  $^{12}\text{B}$
  - **Long-lived: xenon spallation**

# Long-lived isotope production from carbon spallation



# Long-lived isotope production from xenon spallation

Muon spallation on xenon results in long-lived isotopes  $T_{1/2} \sim (\text{sec - months})$



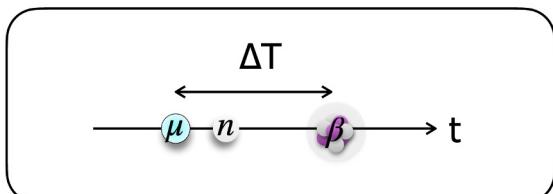
Neutron tagging

**$74.5 \pm 0.4\%$**

LL tagging

**$47.1 \pm 8.7\%$**

Deadtime **9%**

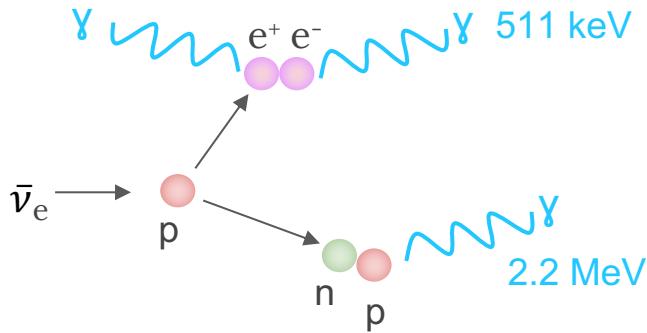
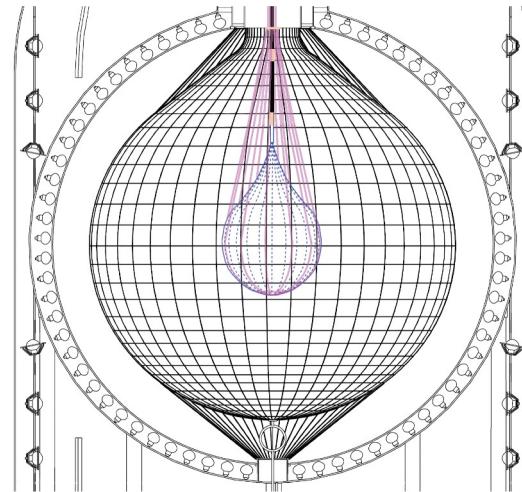


Likelihood based on number of neutrons (ENN), distance ( $dR_{\text{nearest}}$ ) and time ( $dT < 4.6$  days)

## Data selection criteria and fit method

Events selected within analysis volume ( $r < 2.5\text{m}$ ):

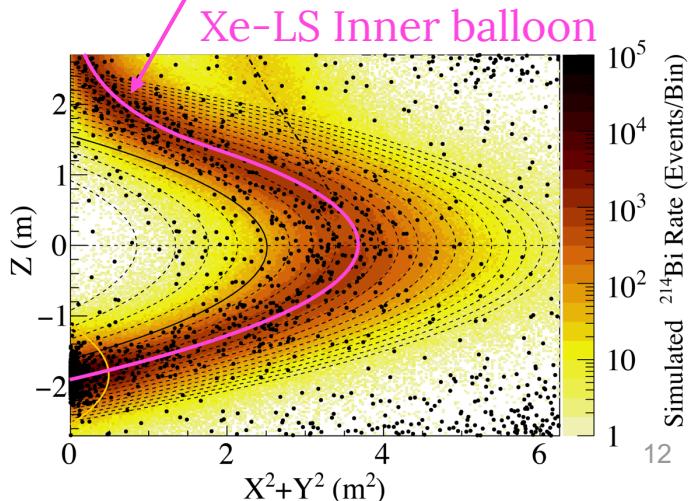
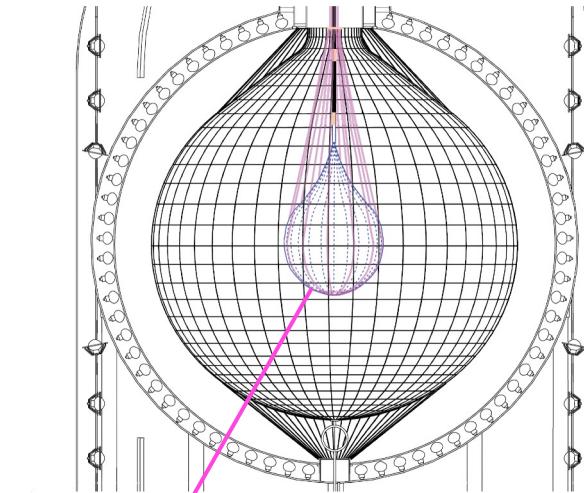
- Muons & 2ms after muons rejected
- $\bar{\nu}$  rejected with inverse- $\beta$  decay coincidence cut
- Short-lived spallation coincidence & shower cut



# Data selection criteria and fit method

Events selected within analysis volume ( $r < 2.5\text{m}$ ):

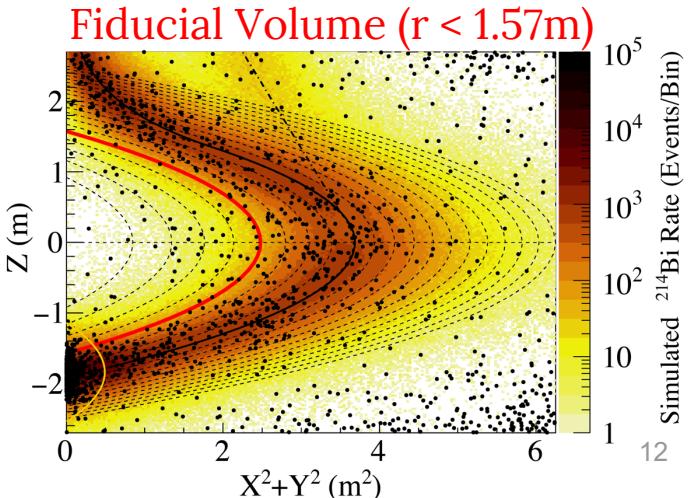
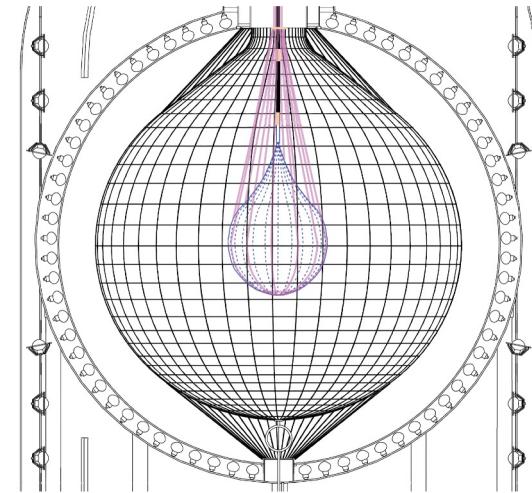
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- RI rejected with BiPo coincidence and FV



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# Data selection criteria and fit method

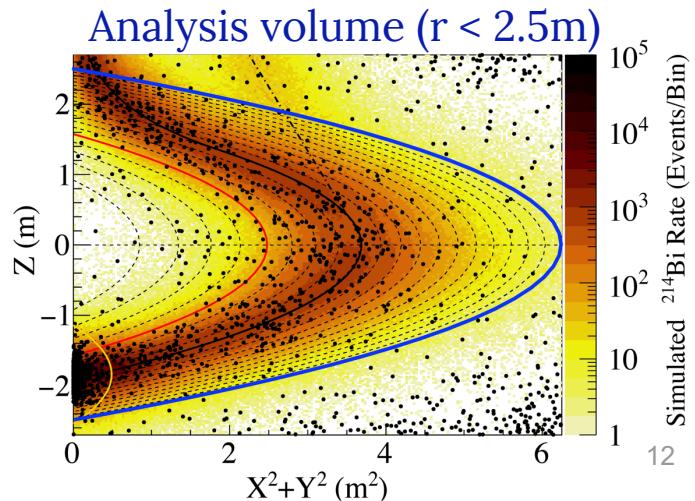
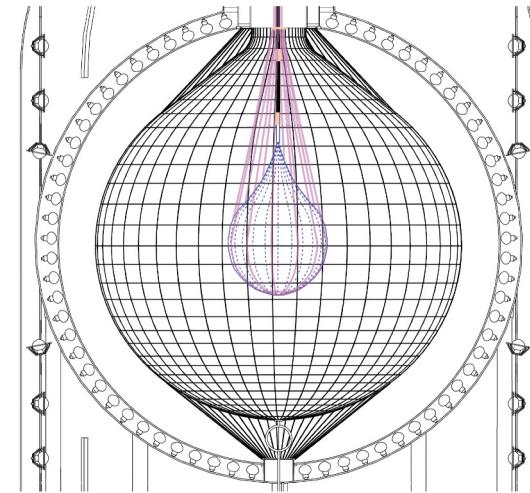
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- RI rejected with BiPo coincidence and FV

Fitting procedure:

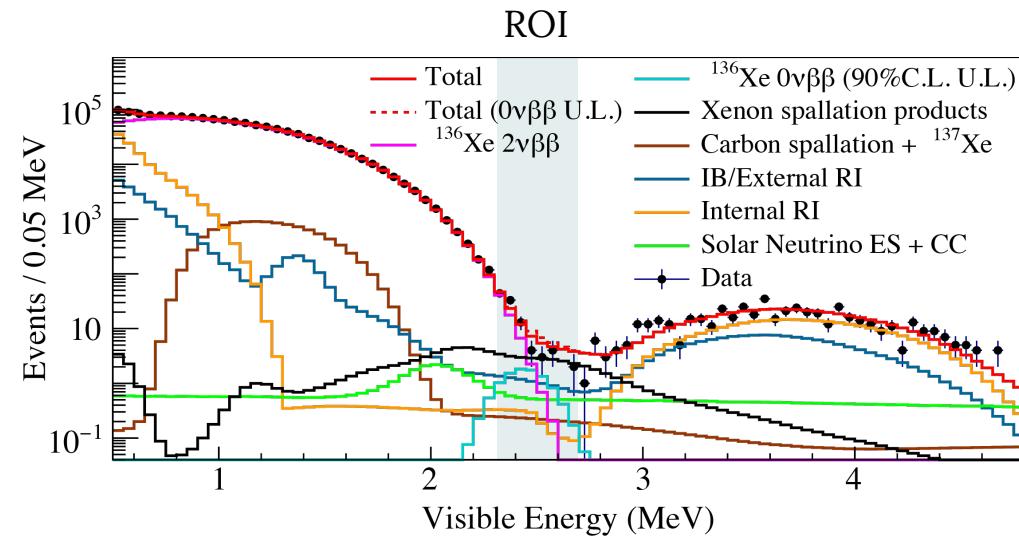
- 40 equal volume bins within  $r < 2.5\text{m}$
- Energy bins  $\in (0.5\text{--}4.8) \text{ MeV} | \text{ROI } (2.35\text{--}2.70) \text{ MeV}$

Simultaneous fit of Long-Lived likelihood tagged (LD) data and  $0\nu\beta\beta$  selected events (SD) to constrain LD rate



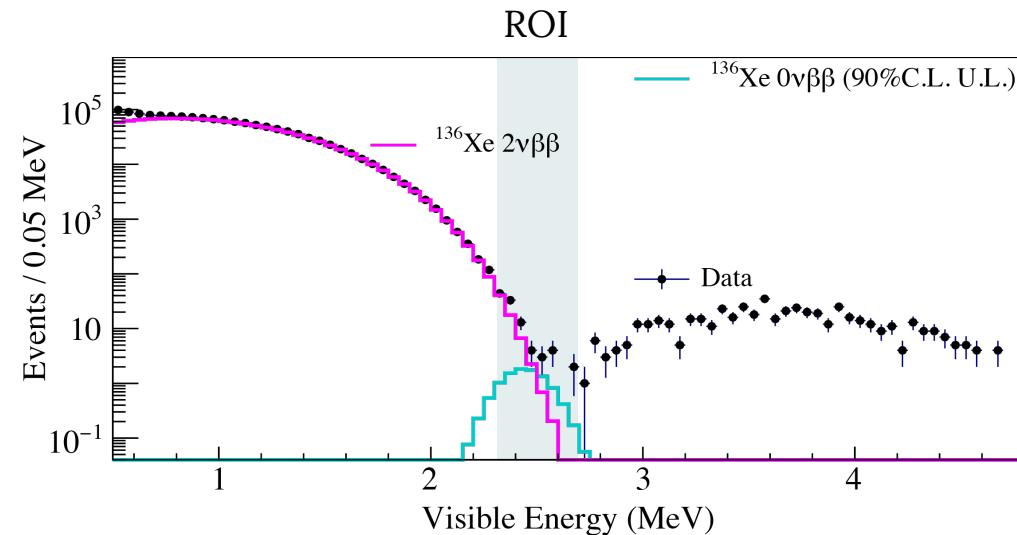
# Most stringent limit on the half-life of $0\nu\beta\beta$

➤ Livetime 1131 days | Exposure 2.1 ton yr  $^{136}\text{Xe}$



# Most stringent limit on the half-life of $0\nu\beta\beta$

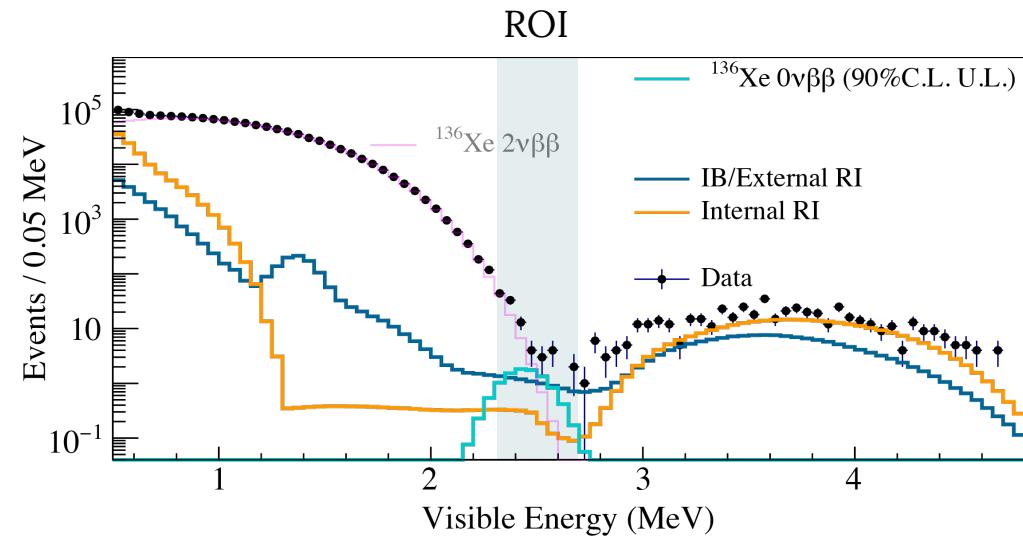
- Livetime 1131 days | Exposure 2.1 ton yr  $^{136}\text{Xe}$



- $2\nu\beta\beta$  decay

# Most stringent limit on the half-life of $0\nu\beta\beta$

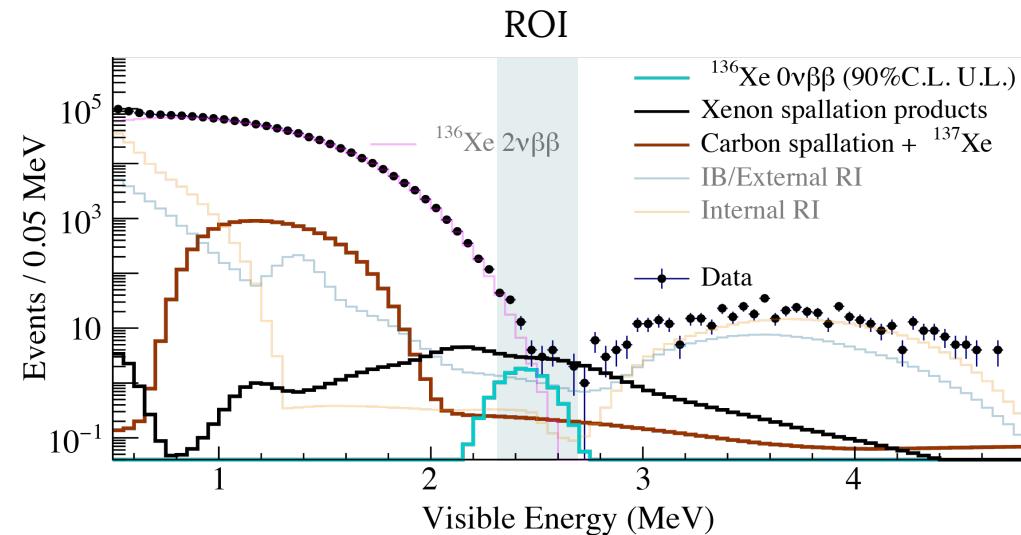
- Livetime 1131 days | Exposure 2.1 ton yr  $^{136}\text{Xe}$



- $2\nu\beta\beta$  decay
- Radioactive contamination

# Most stringent limit on the half-life of $0\nu\beta\beta$

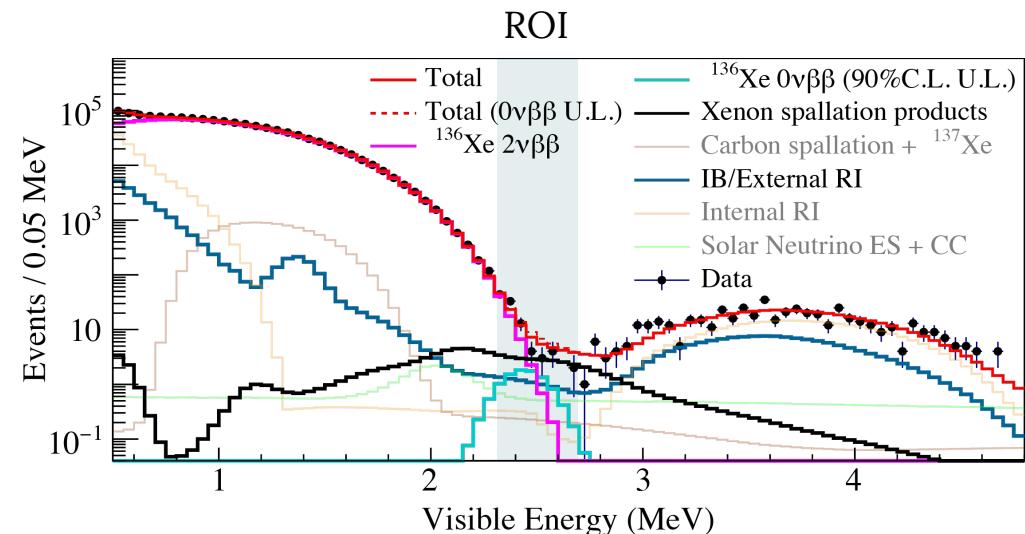
- Livetime 1131 days | Exposure 2.1 ton yr  $^{136}\text{Xe}$



- $2\nu\beta\beta$  decay
- Radioactive contamination
- Cosmogenic spallation products

# Most stringent limit on the half-life of $0\nu\beta\beta$

- Livetime 1131 days | Exposure 2.1 ton yr  $^{136}\text{Xe}$



- $2\nu\beta\beta$  decay
- Radioactive contamination
- Cosmogenic spallation products

## Fitting results (90% C.L.):

Before:  $T_{1/2}^{0\nu} > 2.3 \times 10^{26}$  yr

Latest:  $T_{1/2}^{0\nu} > 3.8 \times 10^{26}$  yr

# Implications for the neutrino mass ordering

$$(T_{1/2}^{0\nu})^{-1} = G_{0\nu}(Q, Z) |M_{0\nu}|^2 \langle m_{\beta\beta} \rangle^2 \longrightarrow \langle m_{\beta\beta} \rangle = \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$$

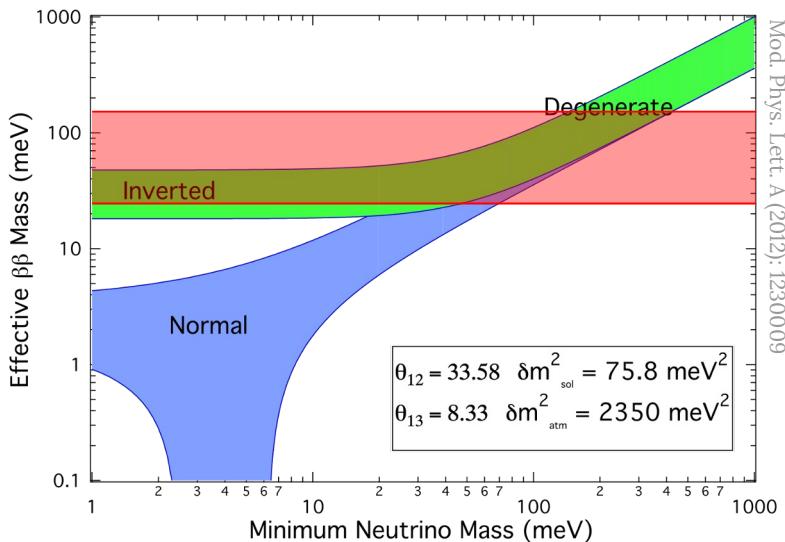
$$T_{1/2}^{0\nu} > 3.8 \times 10^{26} \text{ yr}$$

Nuclear Matrix Element

$$\langle m_{\beta\beta} \rangle < (28 - 122) \text{ meV}$$

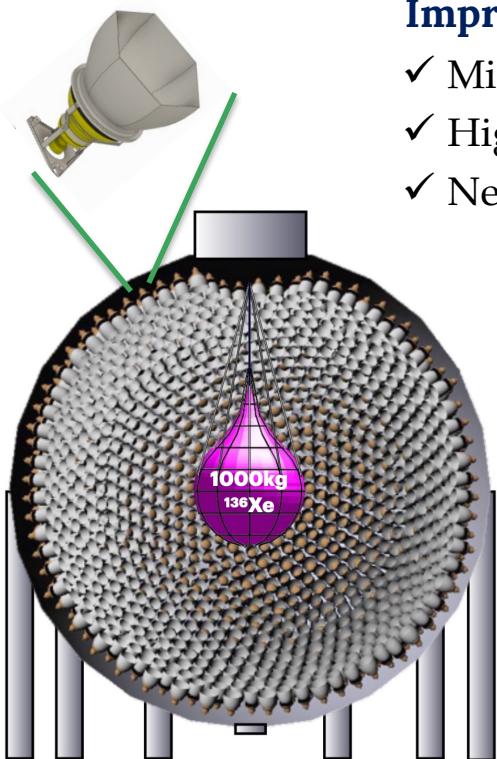
$$m_{\nu 1} > m_{\nu 3}$$

$$m_{\nu 3} > m_{\nu 1}$$



# Upgrade to KamLAND2-Zen

Last physics run on August 27<sup>th</sup> 2024 → dismantling of detector started



## Improved energy resolution

- ✓ Mirrors around PMTs (x1.8)
- ✓ Higher QE PMTs (x1.9)
- ✓ New liquid scintillator (x1.4)

4% → 2% at 2.6 MeV

After 5 year:

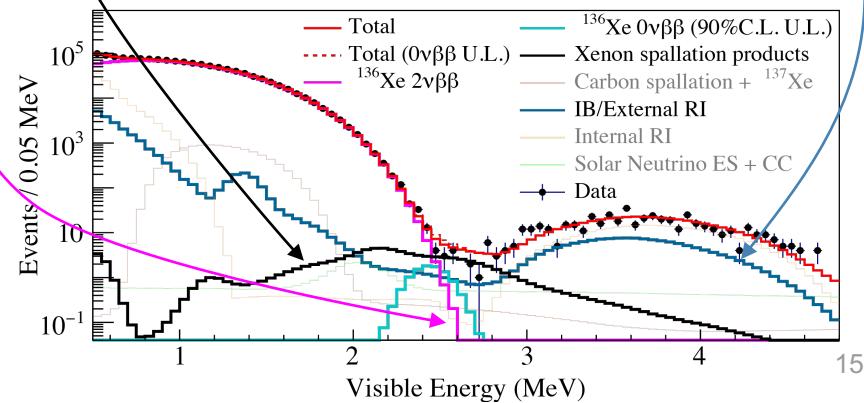
$$T_{1/2}^{0\nu} > 2 \times 10^{27} \text{ yr}$$

## Scintillation inner balloon (PEN)

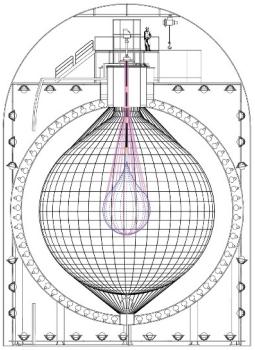
→ Larger fiducial volume

## Dead-time free electronics

→ Higher neutron tagging efficiency



# KamLAND2-Zen construction timeline



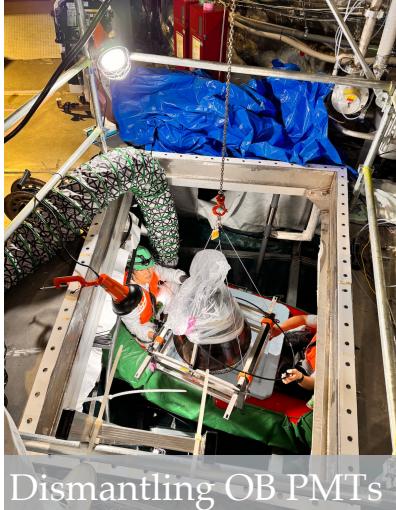
## KamLAND dismantling

2024

- ✓ Xe & LS extraction
  - LS extraction
  - Dismantling OB & PMTs



Water tank

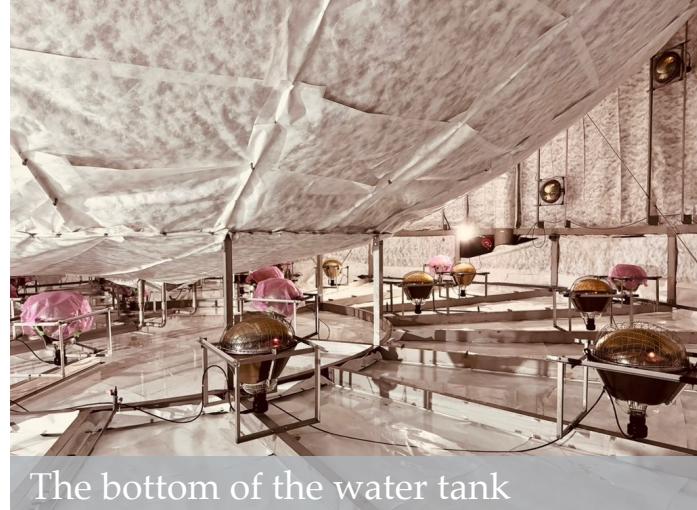


Dismantling OB PMTs

## KamLAND2 construction

2027

- PMT & mirror installation
- Cleaner OB | LS | new electronics
- Calibration system



The bottom of the water tank

# Future prospects

---

KamLAND-Zen is searching for  $0\nu\beta\beta$  with  $^{136}\text{Xe}$  dissolved in liquid scintillator

- Current most stringent limit of is  $T_{1/2} > 3.8 \times 10^{26} \text{ yr} \rightarrow m_{\beta\beta} < (28 - 122)\text{meV}$

The detector upgrade of KamLAND to KamLAND2 has started

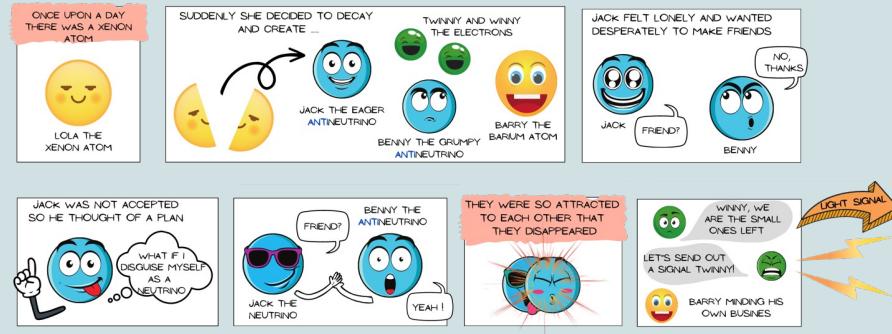
- Scintillating inner balloon
- Improved energy resolution
- New electronics for enhanced tagging efficiency

KamLAND2-Zen aims to cover the IO region | target sensitivity  $\langle m_{\beta\beta} \rangle = 20 \text{ meV}$

**KamLAND2 dismantling & construction ongoing → expected launch in 2027!**

# Thank you for your attention!

Follow me for fun videos  
& comics on KamLAND



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Kelly Weerman  
Experimental particle physics PhD student working on the KamLAND detector in Japan

← physicswithkelly ⚙️ :

17 posts 5.792 followers 43 following

Kelly Weerman  
Scientist  
Experimental particle physics PhD student working on the KamLAND detector in Japan

we do that by opening this valve

why is there more matter than antimatter

these electronics can get really warm

so when a neutrino, which is a super light particle, flies through our detector

This is actually the top part of the inner detector, or around

# **Backup slides**

# KamLAND-Zen $0\nu\beta\beta$ half-life and mass limits

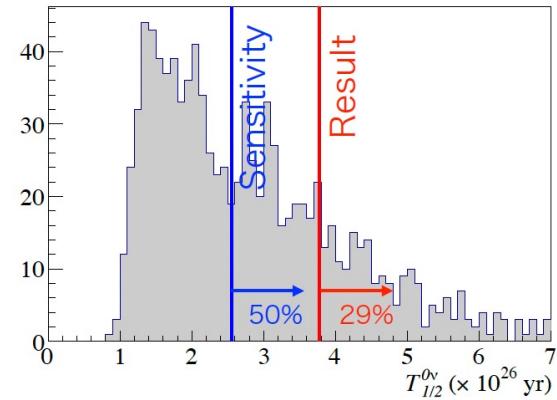
Majorana mass expectation from the KamLAND-Zen half-life limit using different nuclear matrix elements

	$M^{0\nu}$	$\langle m_{\beta\beta} \rangle$ (meV)
Shell model	2.28, 2.45	59.4, 55.3
	1.63, 1.76	83.1, 77.0
	2.39	56.7
QRPA	1.55	87.4
	2.91	46.6
	2.71	50.0
	1.11, 1.18	122, 115
	3.38	40.1
EDF theory	4.20	32.3
	4.77	28.4
	4.24	32.0
IBM	3.25	41.7
	3.40	39.9

Confidence limits

**KamLAND-Zen 800 only:**  
 $T_{1/2} > 3.4 \times 10^{26}$  yr

**Combined Zen 400 + 800:**  
 $T_{1/2} > 3.8 \times 10^{26}$  yr



	Dataset	Exposure	$T_{1/2}$ (90% C.L.)
KamLAND-Zen 400	Phase I	85.9 kg yr	$1.9 \times 10^{25}$ yr
	+ Phase II	504 kg yr	$1.07 \times 10^{26}$ yr
KamLAND-Zen 800	First dataset	970 kg yr	$2.3 \times 10^{26}$ yr
	Complete dataset	2097 kg yr	$3.8 \times 10^{26}$ yr

# Majorana neutrino mass prediction

**Dark shaded regions:** predictions based on best-fit values of neutrino oscillation parameters

**Light shaded regions:**  $3\sigma$  ranges from oscillation parameter uncertainties.

## NME calculations:

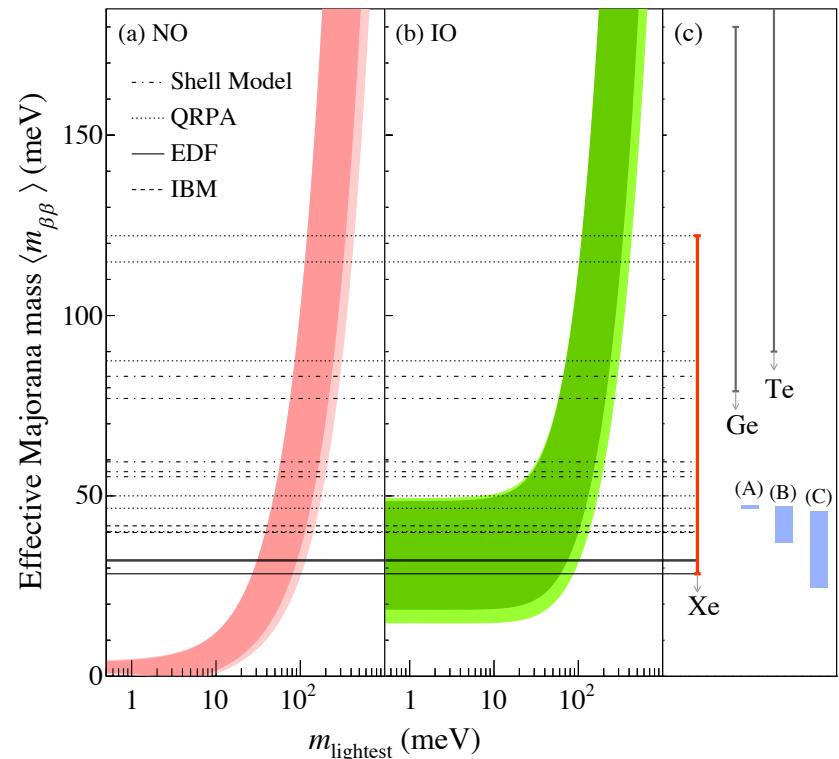
Shell model → dot-dashed lines

QRPA → dotted lines

EDF theory → solid lines

IBM → dashed lines

**Three theoretical predictions** in IO region: A, B, C



# KamLAND-Zen $0\nu\beta\beta$ confidence limits

Frequentist confidence limit (Wilks'):

$$T_{1/2}^{0\nu\beta\beta} > 3.4 \times 10^{26} \text{yr (90 \% C.L.)}$$

Frequentist Feldman-Cousins calculation result:

$$T_{1/2}^{0\nu\beta\beta} > 4.3 \times 10^{26} \text{yr (90 \% C.L.)}$$

Combined KLZ 400 + 800 Analysis (Wilks')

$$T_{1/2}^{0\nu\beta\beta} > 3.8 \times 10^{26} \text{yr (90 \% C.L.)}$$

Bayesian result:

$$T_{1/2}^{0\nu\beta\beta} > 3.4 \times 10^{26} \text{yr (90 \% C.I.)}$$

# Experimental design criteria for $0\nu\beta\beta$

Direct searches: kinematic parameters of the two electrons

→ Total energy and individual electron paths

$$T_{1/2}^{0\nu} = \ln 2 \frac{N_A}{W} \left( \frac{a \cdot \epsilon \cdot M}{N_{\text{obs}}} \right) t \propto \begin{cases} a\epsilon \cdot Mt \\ a\epsilon \sqrt{\frac{Mt}{N_{\text{bkg}} \cdot \Delta E}} \end{cases}$$

Detector and isotope choice depending on:

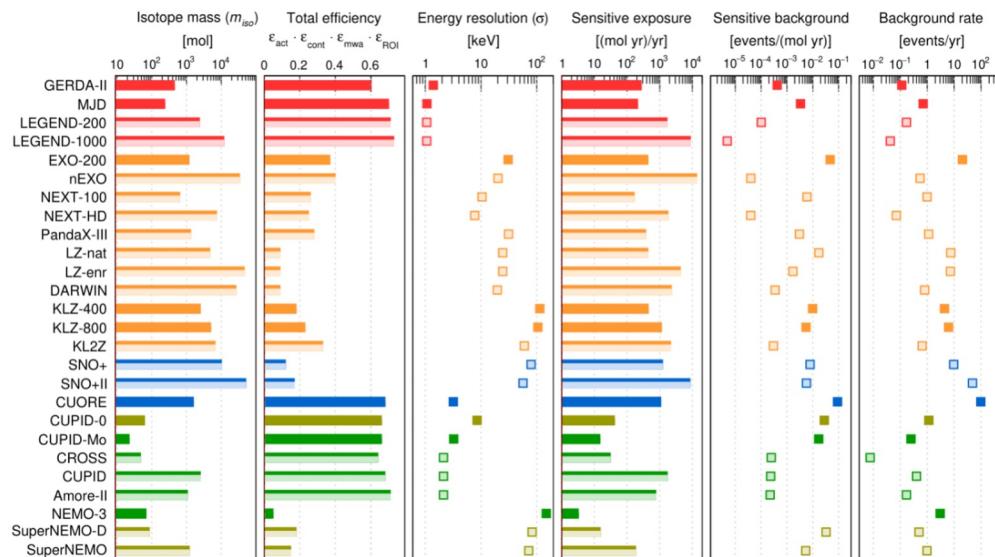
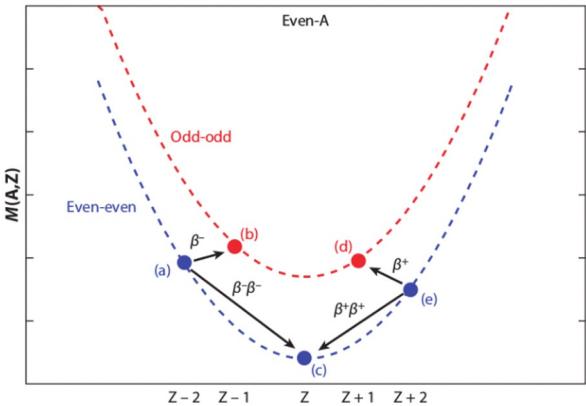
- High isotopic abundance  $a$
- Deployment in large quantity  $M$
- High-resolution detector  $\Delta E$
- Low-background conditions  $N_{\text{bkg}}$

# Different isotope choices for detecting $0\nu\beta\beta$

Isotope	$G^{0\nu}$ ( $10^{-14} \text{ y}^{-1}$ )	$Q_{\beta\beta}$ (keV)	Nat. ab. (%)
$^{48}\text{Ca}$	6.35	4273.7	0.187
$^{76}\text{Ge}$	0.623	2039.1	7.8
$^{82}\text{Se}$	2.70	2995.5	9.2
$^{96}\text{Zr}$	5.63	3347.7	2.8
$^{100}\text{Mo}$	4.36	3035.0	9.6
$^{110}\text{Pd}$	1.40	2004.0	11.8
$^{116}\text{Cd}$	4.62	2809.1	7.6
$^{124}\text{Sn}$	2.55	2287.7	5.6
$^{130}\text{Te}$	4.09	2530.3	34.5
$^{136}\text{Xe}$	4.31	2461.9	8.9
$^{150}\text{Nd}$	19.2	3367.3	5.6

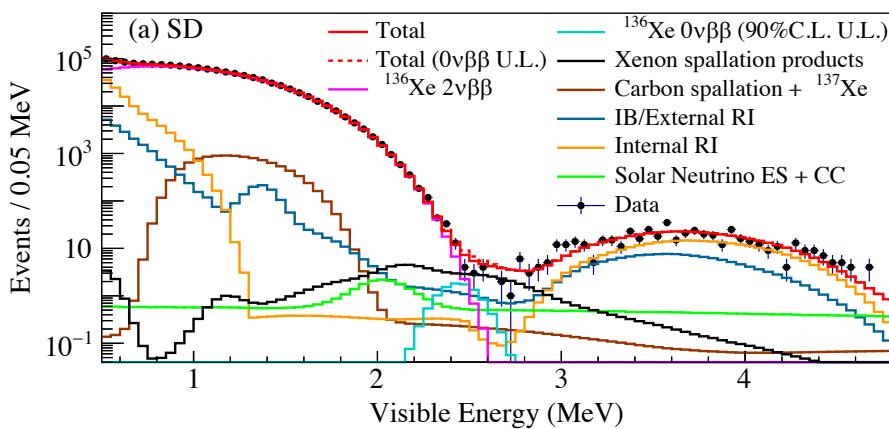
Isotope	Abundance (%)	$Q_{\beta\beta}$ (MeV)	$G^{2\nu} (10^{-18} \text{ year}^{-1})$
$^{48}\text{Ca}$	0.187	4.263	15.6
$^{76}\text{Ge}$	7.8	2.039	0.0482
$^{82}\text{Se}$	9.2	2.998	1.60
$^{96}\text{Zr}$	2.8	3.348	7.83
$^{100}\text{Mo}$	9.6	3.035	4.13
$^{116}\text{Cd}$	7.6	2.813	3.18
$^{130}\text{Te}$	34.08	2.527	1.53
$^{136}\text{Xe}$	8.9	2.459	1.43
$^{150}\text{Nd}$	5.6	3.371	36.4

$^{76}\text{Ge}$   
 $^{136}\text{Xe}$   
 $^{130}\text{Te}$   
 $^{100}\text{Mo}$   
 $^{82}\text{Se}$

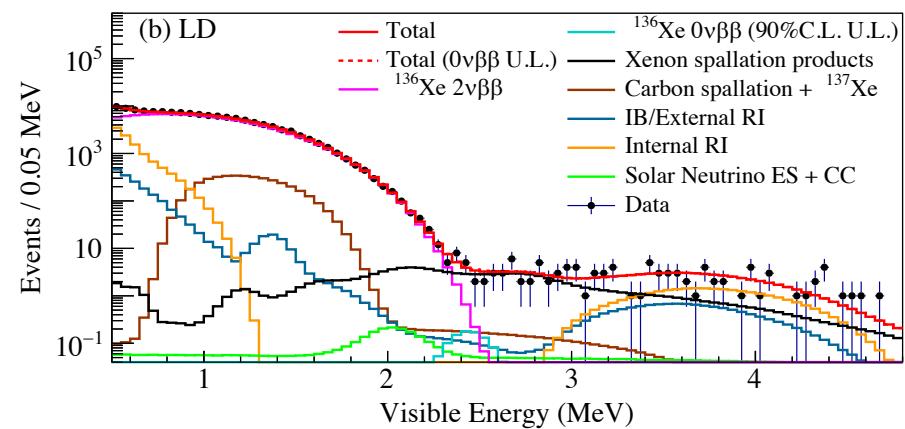


# KamLAND-Zen Singles- and LL-spectrum

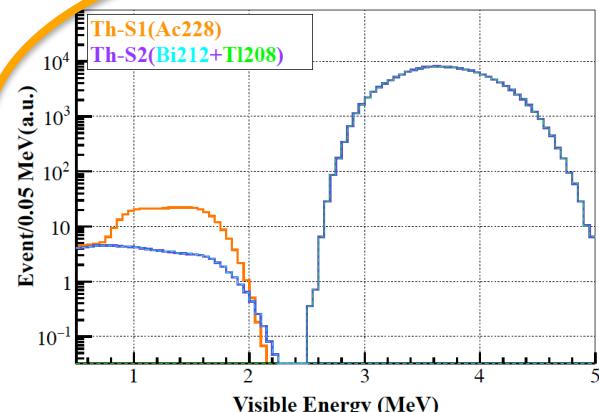
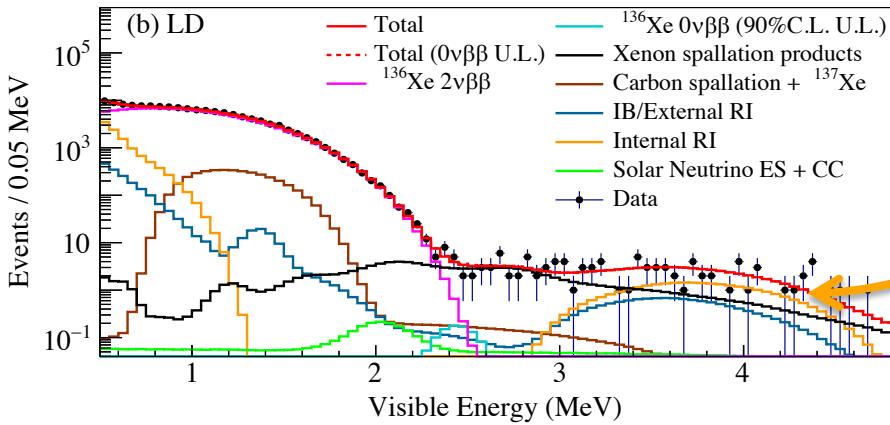
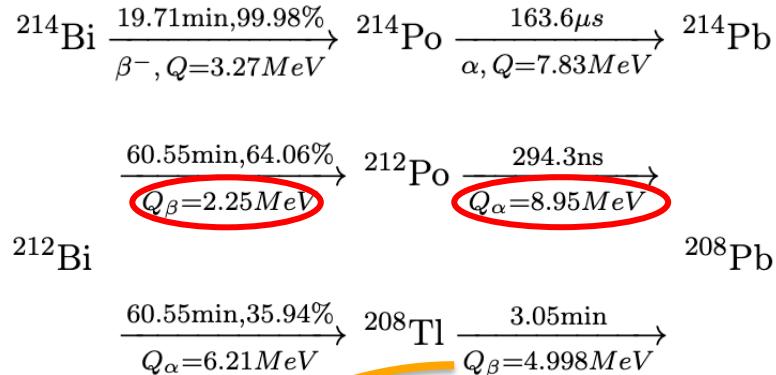
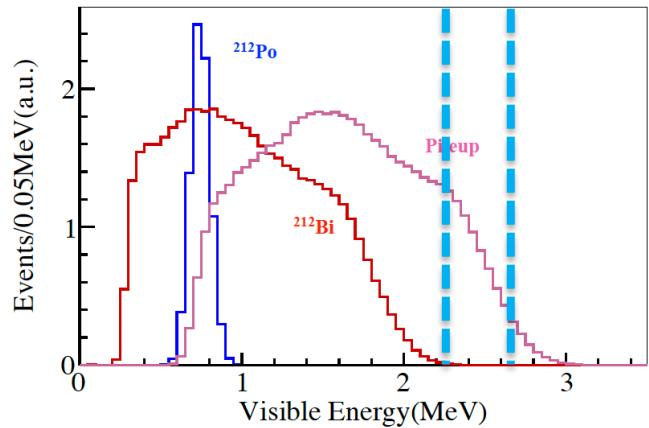
Singles data:  $0\nu\beta\beta$  candidates



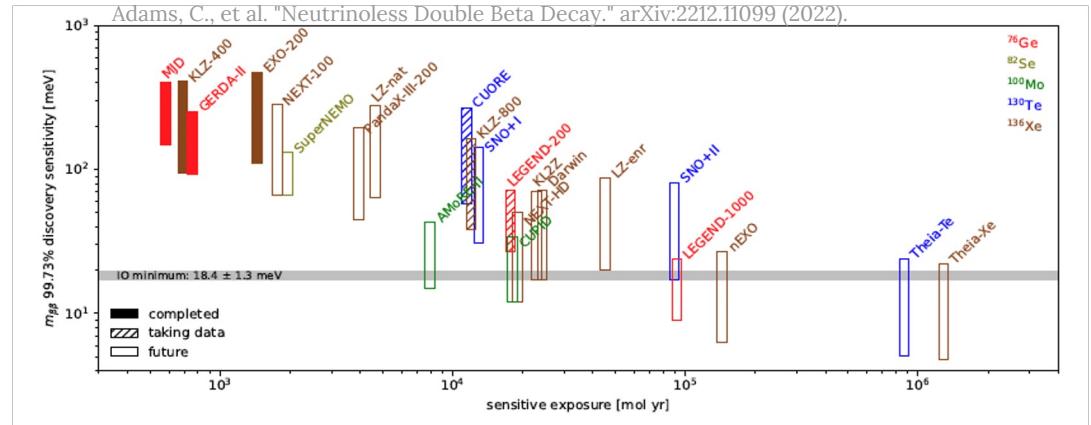
Long-lived data: LL-candidates | 10% of SD



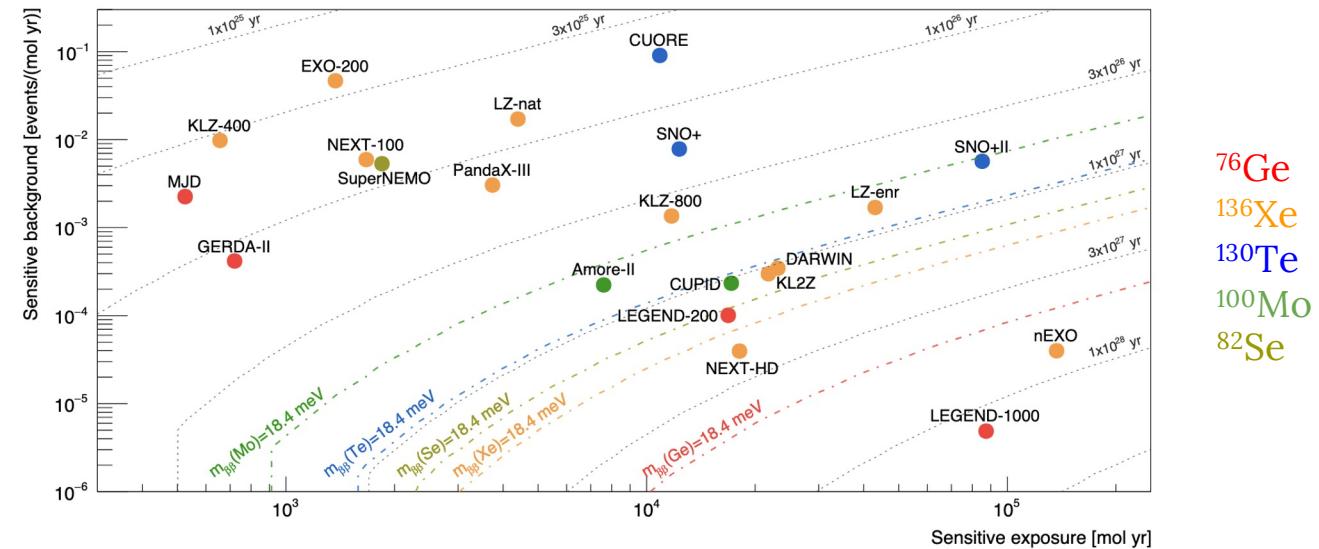
# Radioactive background: $^{232}\text{Th}$ decay series



# $0\nu\beta\beta$ detector limits

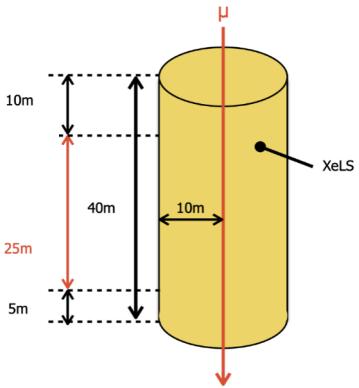


Agostini, Matteo, et al. Rev. Mod. Phys. 95, 025002 (2023)



# Xenon spallation FLUKA simulations

FLUKA simulation geometry setup  
 → 9yr detector livetime



FLUKA simulated muon spallation isotopes in KamLAND-Zen Xe-LS

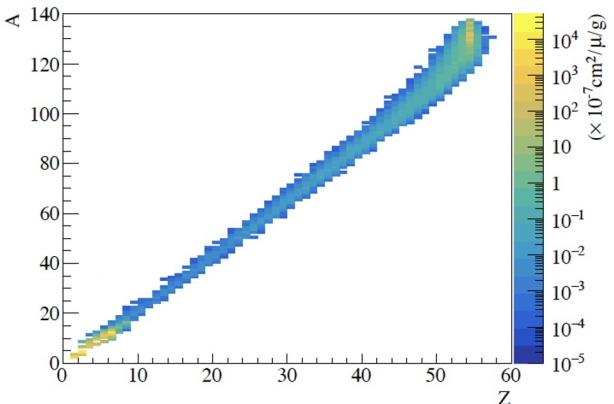


TABLE IX. Simulated production rate of dominant isotopes in  $2.35 \leq E \leq 2.70$  MeV in Xe-LS.

	$\tau_{1/2}$ (s)	$Q$ (MeV)	(kton day) $^{-1}$
			ROI
$^{88}\text{Y}$	$9.212 \times 10^6$	$3.62 (\text{EC}/\beta^+\gamma)$	0.110 0.136
$^{90m1}\text{Zr}$	$8.092 \times 10^{-1}$	$2.31 (\text{IT})$	0.012 0.093
$^{90}\text{Nb}$	$5.256 \times 10^4$	$6.11 (\text{EC}/\beta^+\gamma)$	0.024 0.095
$^{96}\text{Tc}$	$3.698 \times 10^5$	$2.97 (\text{EC}/\beta^+\gamma)$	0.012 0.059
$^{98}\text{Rh}$	$5.232 \times 10^2$	$5.06 (\text{EC}/\beta^+\gamma)$	0.011 0.076
$^{100}\text{Rh}$	$7.488 \times 10^4$	$3.63 (\text{EC}/\beta^+\gamma)$	0.088 0.234
$^{104}\text{Ag}$	$4.152 \times 10^3$	$4.28 (\text{EC}/\beta^+\gamma)$	0.012 0.160
$^{104m1}\text{Ag}$	$2.010 \times 10^3$	$4.28 (\text{EC}/\beta^+\gamma)$	0.018 0.111
$^{107}\text{In}$	$1.944 \times 10^3$	$3.43 (\text{EC}/\beta^+\gamma)$	0.019 0.135
$^{108}\text{In}$	$3.480 \times 10^3$	$5.16 (\text{EC}/\beta^+\gamma)$	0.089 0.194
$^{110}\text{In}$	$1.771 \times 10^4$	$3.89 (\text{EC}/\beta^+\gamma)$	0.053 0.236
$^{110m1}\text{In}$	$4.146 \times 10^3$	$3.89 (\text{EC}/\beta^+\gamma)$	0.066 0.351
$^{109}\text{Sn}$	$1.080 \times 10^3$	$3.85 (\text{EC}/\beta^+\gamma)$	0.027 0.122
$^{113}\text{Sb}$	$4.002 \times 10^2$	$3.92 (\text{EC}/\beta^+\gamma)$	0.036 0.231
$^{114}\text{Sb}$	$2.094 \times 10^2$	$5.88 (\text{EC}/\beta^+\gamma)$	0.020 0.297
$^{115}\text{Sb}$	$1.926 \times 10^3$	$3.03 (\text{EC}/\beta^+\gamma)$	0.031 0.839
$^{116}\text{Sb}$	$9.480 \times 10^2$	$4.71 (\text{EC}/\beta^+\gamma)$	0.071 0.939
$^{118}\text{Sb}$	$2.160 \times 10^2$	$3.66 (\text{EC}/\beta^+\gamma)$	0.165 1.288
$^{124}\text{Sb}$	$5.201 \times 10^6$	$2.90 (\text{EC}/\beta^-\gamma)$	0.016 0.054
$^{115}\text{Te}$	$3.480 \times 10^2$	$4.64 (\text{EC}/\beta^+\gamma)$	0.012 0.124
$^{117}\text{Te}$	$3.720 \times 10^3$	$3.54 (\text{EC}/\beta^+\gamma)$	0.052 0.594
$^{119}\text{I}$	$1.146 \times 10^3$	$3.51 (\text{EC}/\beta^+\gamma)$	0.053 0.533
$^{120}\text{I}$	$4.896 \times 10^3$	$5.62 (\text{EC}/\beta^+\gamma)$	0.091 0.953
$^{122}\text{I}$	$2.178 \times 10^2$	$4.23 (\text{EC}/\beta^+\gamma)$	0.289 1.965
$^{124}\text{I}$	$3.608 \times 10^5$	$3.16 (\text{EC}/\beta^+\gamma)$	0.190 1.654
$^{130}\text{I}$	$4.450 \times 10^4$	$2.95 (\beta^-\gamma)$	0.195 1.188
$^{132}\text{I}$	$8.262 \times 10^3$	$3.58 (\beta^-\gamma)$	0.148 0.427
$^{134}\text{I}$	$3.150 \times 10^3$	$4.18 (\beta^-\gamma)$	0.043 0.183
$^{121}\text{Xe}$	$2.406 \times 10^3$	$3.75 (\text{EC}/\beta^+\gamma)$	0.100 0.540
$^{125}\text{Cs}$	$2.802 \times 10^3$	$3.09 (\text{EC}/\beta^+\gamma)$	0.012 0.266
$^{126}\text{Cs}$	$9.840 \times 10^1$	$4.82 (\text{EC}/\beta^+\gamma)$	0.011 0.080
$^{128}\text{Cs}$	$2.196 \times 10^2$	$3.93 (\text{EC}/\beta^+\gamma)$	0.031 0.229